

Simultaneous photometry and spectroscopy of the *exceptional* asteroid (3200) Phaethon

Huib Henrichs
Amsterdam

Salzburg, 4 May 2019

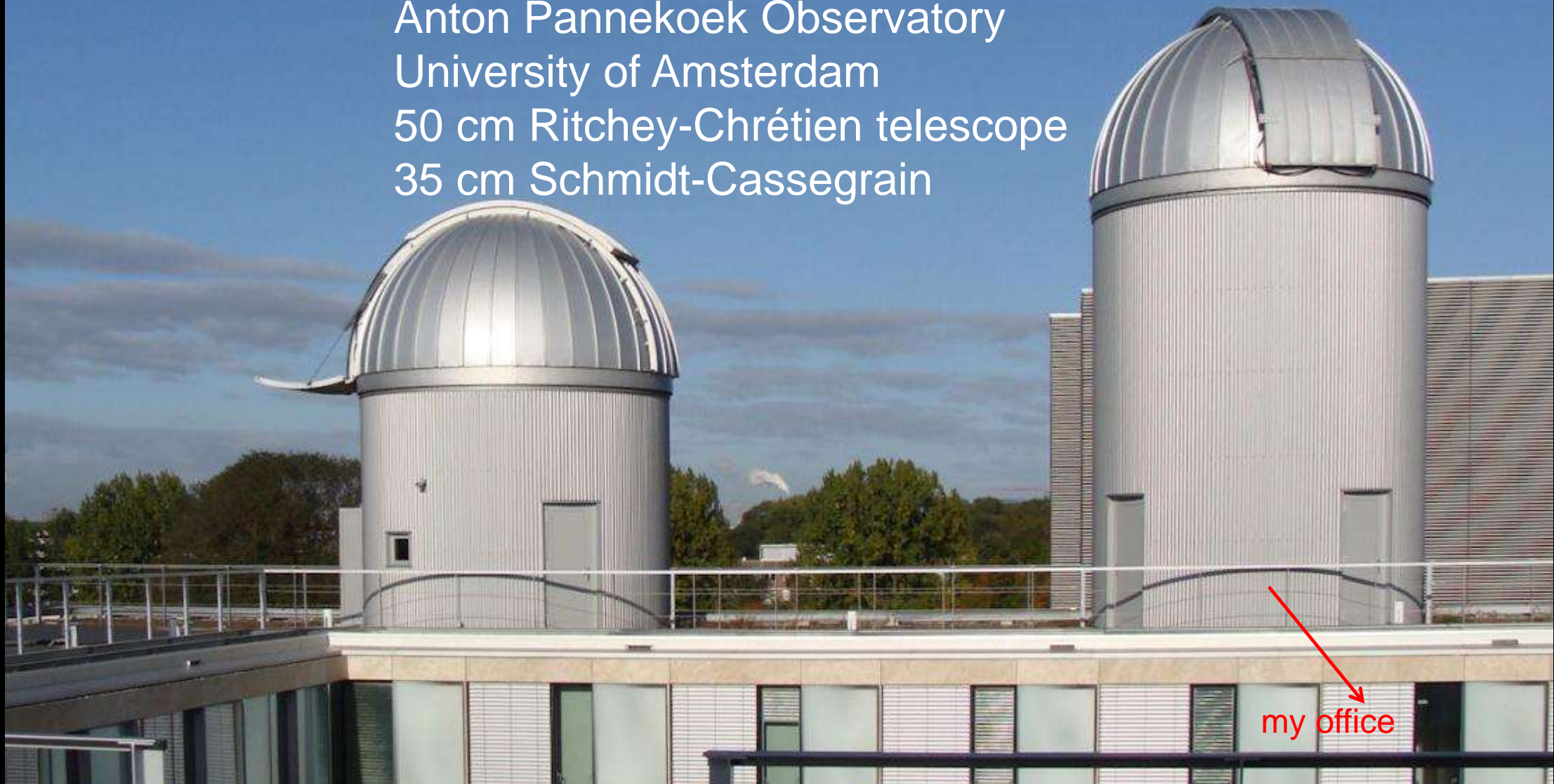
Unique opportunity:

Phaethon closely passed the Earth on December 2017

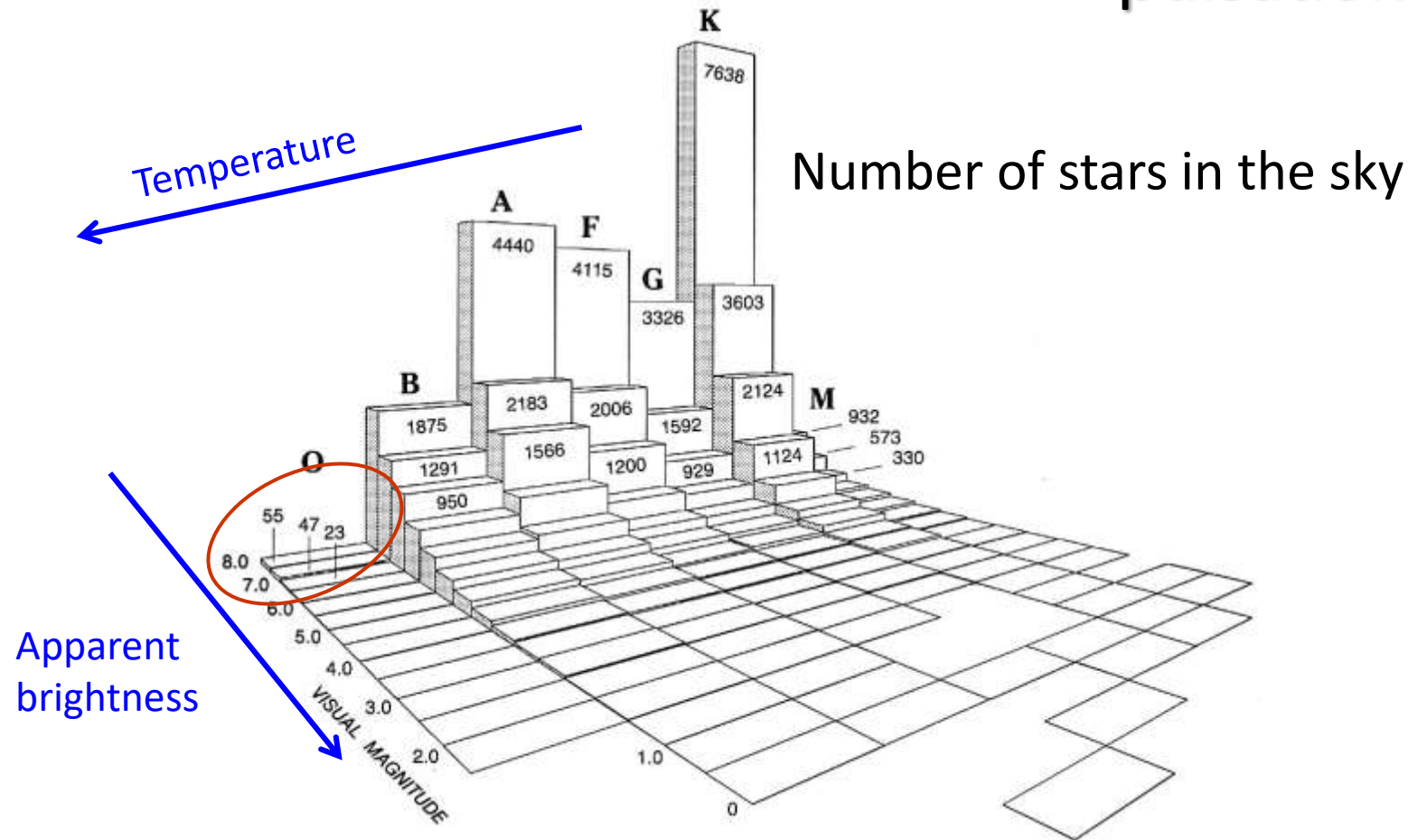
1. Technical part: my remotely controlled observatory in Amsterdam
2. Example spectroscopy projects: supernova, quasars
3. The exceptional Phaethon:
simultaneous spectroscopy and photometry

*“One telescope will be permanently pointed to the Sun,
the other to the rest of the universe”*

Anton Pannekoek Observatory
University of Amsterdam
50 cm Ritchey-Chrétien telescope
35 cm Schmidt-Cassegrain

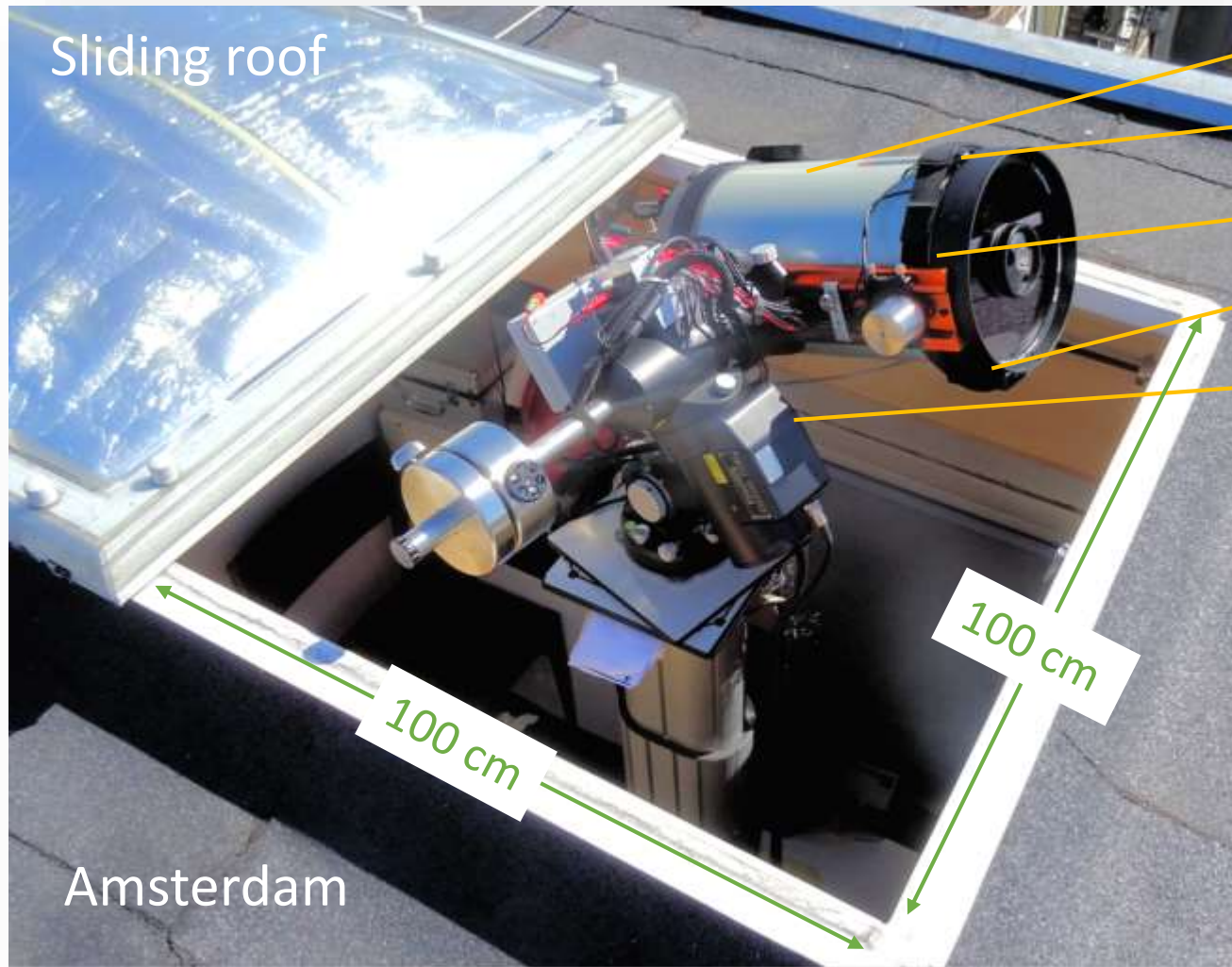


I study massive stars type O and B: magnetic fields
stellar winds
pulsations

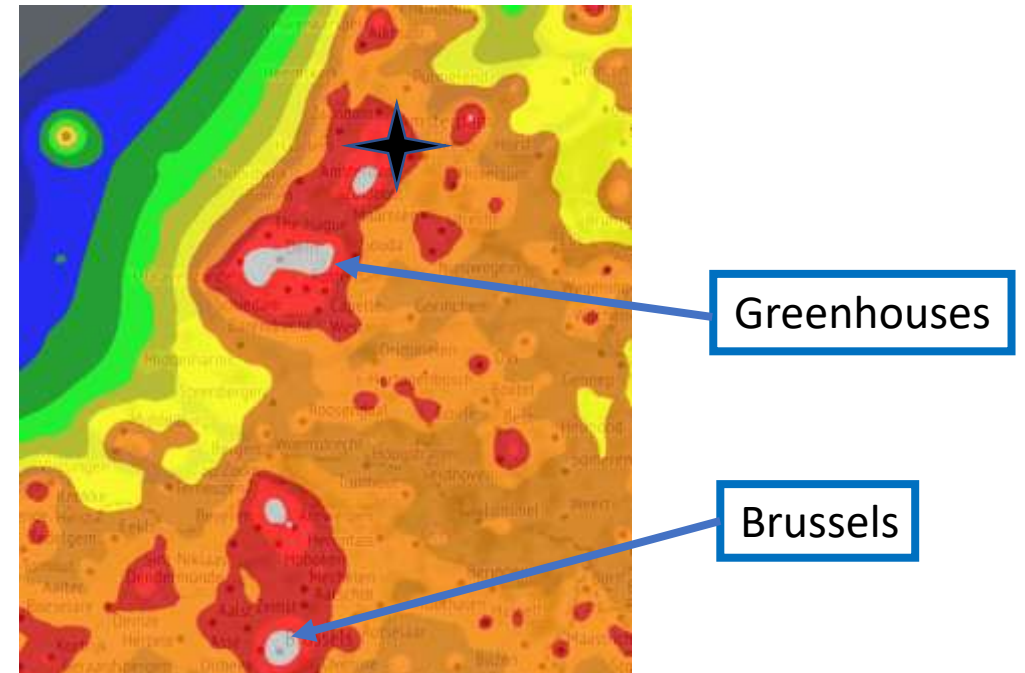


O stars are very rare: ~ 50 visible with the naked eye

1. Rooftop observatory and equipment



- Celestron 9.25" Edge, 23.5 cm diameter
- GPS
- Dew heater
- Unihedron *Sky Quality Meter* (clouds!)
- 10micron GM1000 (absolute encoders)

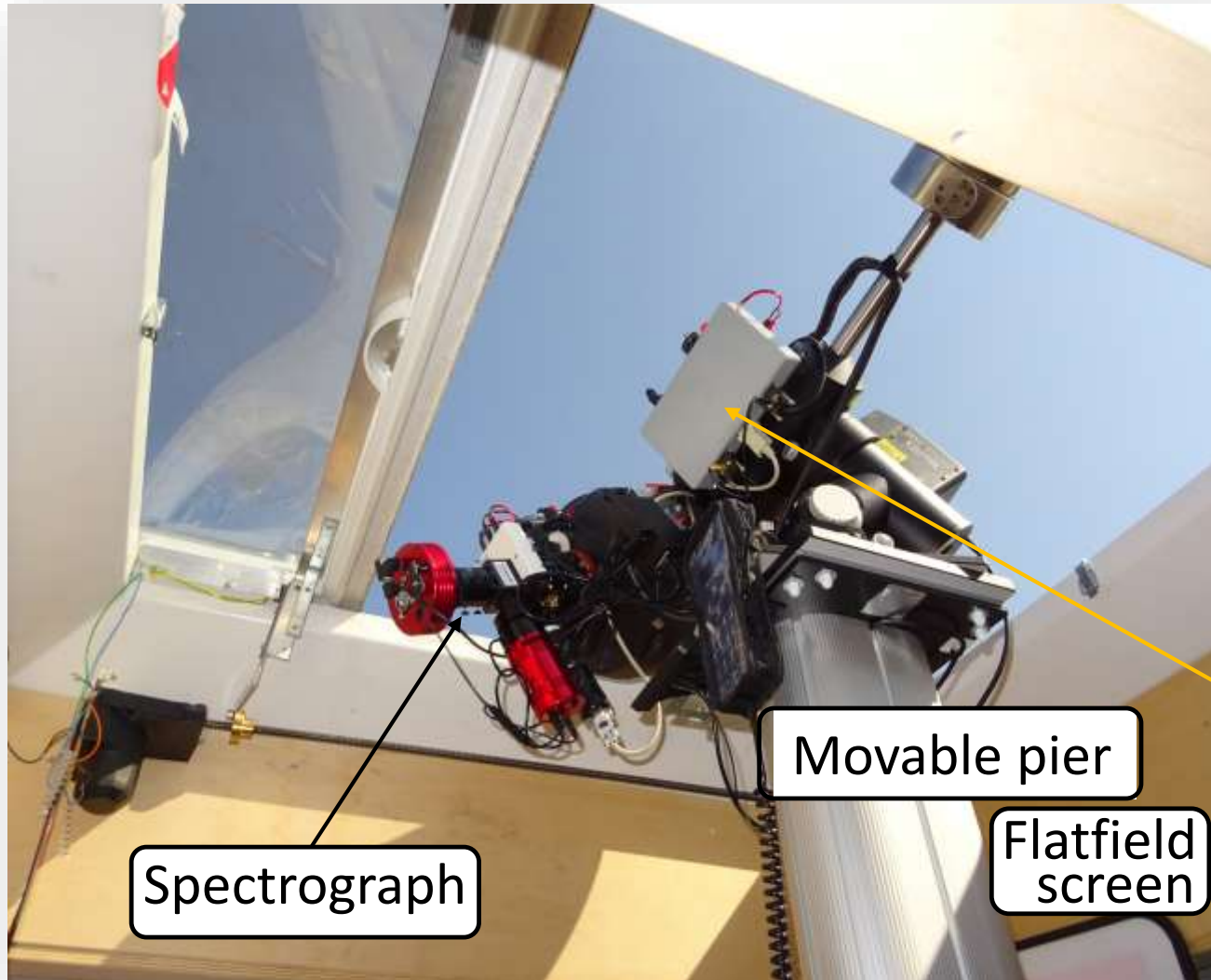




Laptop
with webcam

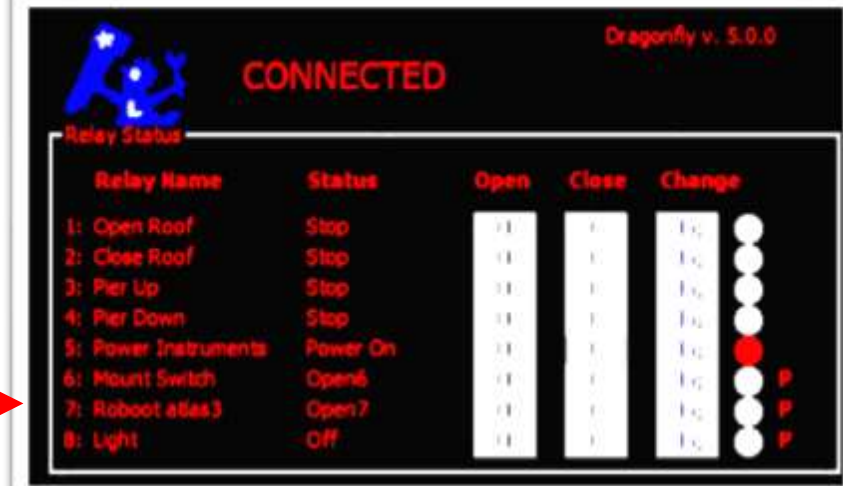
The image shows a custom-built robotic platform housed within a white-painted wooden frame. The platform is constructed from a light-colored material, possibly aluminum or plastic. A laptop is mounted on the left side, with a webcam positioned above it. A large, black, circular cooling fan is mounted on the right side. The fan has a black frame with 'CALSTRON' and '12.500mm' printed on it. The fan blades are black with a green center. A silver, cylindrical component, possibly a motor or a sensor, is mounted on the left side. The entire setup is connected by various wires and cables. The platform is sitting on a grey, textured surface, possibly a concrete floor. The background is a plain, light-colored wall.

Remote view from laptop



Remote control through internet with *Lunatico Dragonfly*:

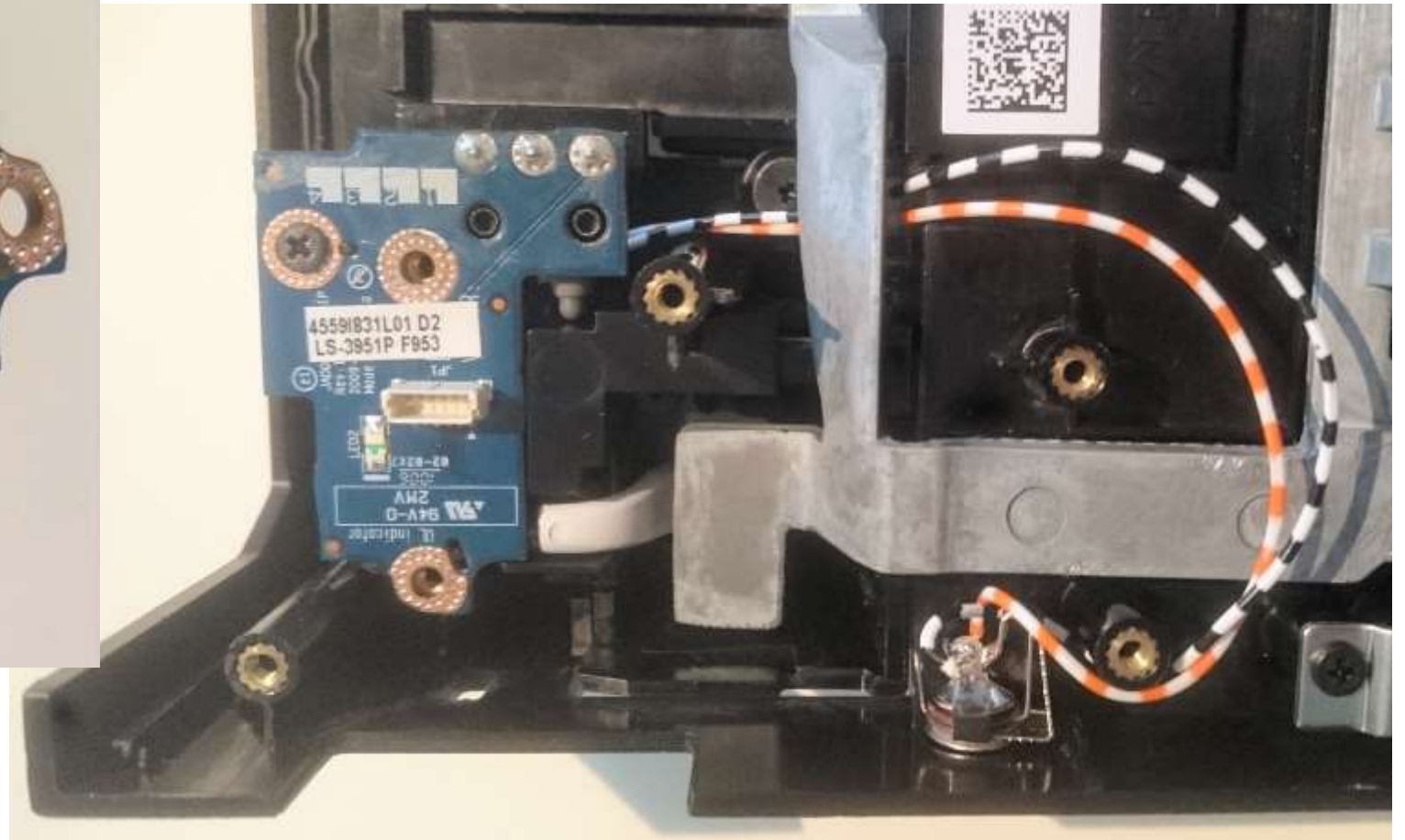
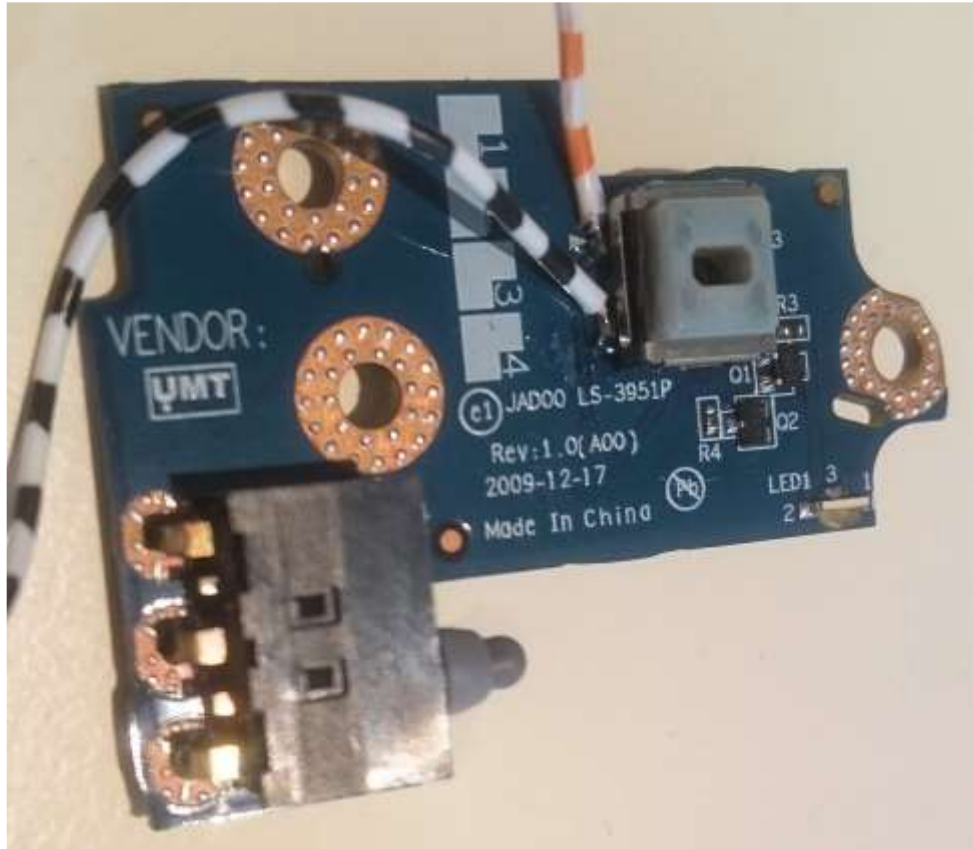
Roof, Pier, Power, Mount, Laptop reboot
through docking station ↗



Instrument control by *Lunatico Platypus*:
focusers + dew heater
+ spectrograph calibration lamps

Software: *Teamviewer*

Laptop remote switch



Projects

➤ Astro Art

Neodymium light-pollution block

Imaging with filters:

- Broadband Nd, RGB (without disturbing moon): galaxies etc.
- Narrowband H α , OIII, SII: nebulae

Fast color: moon/planets

➤ Science

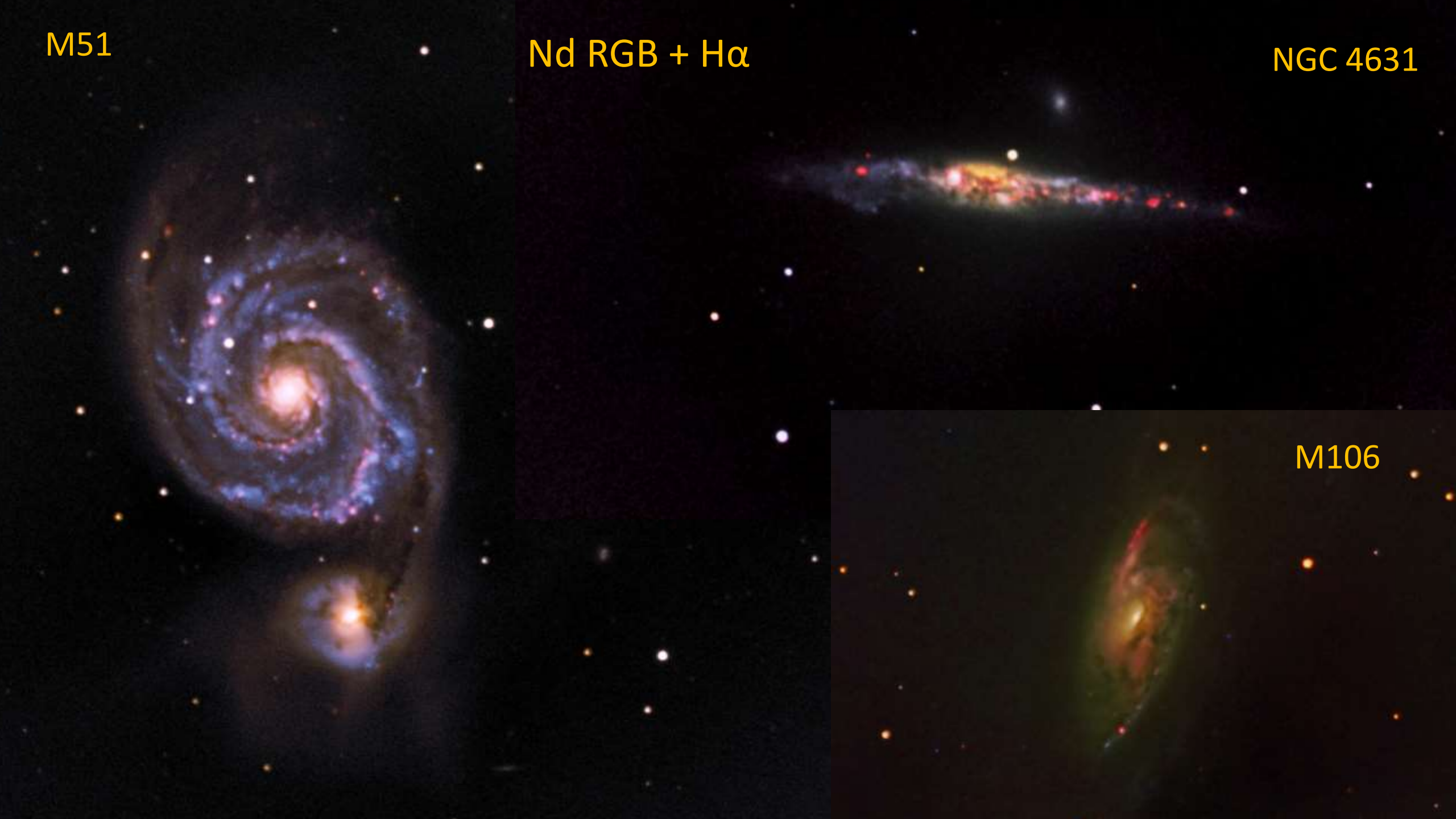
- *Photometry RVB*: supernovae, transients, asteroids
- *Spectroscopy*: quasars, supernovae, stars, asteroids

M51

Nd RGB + H α

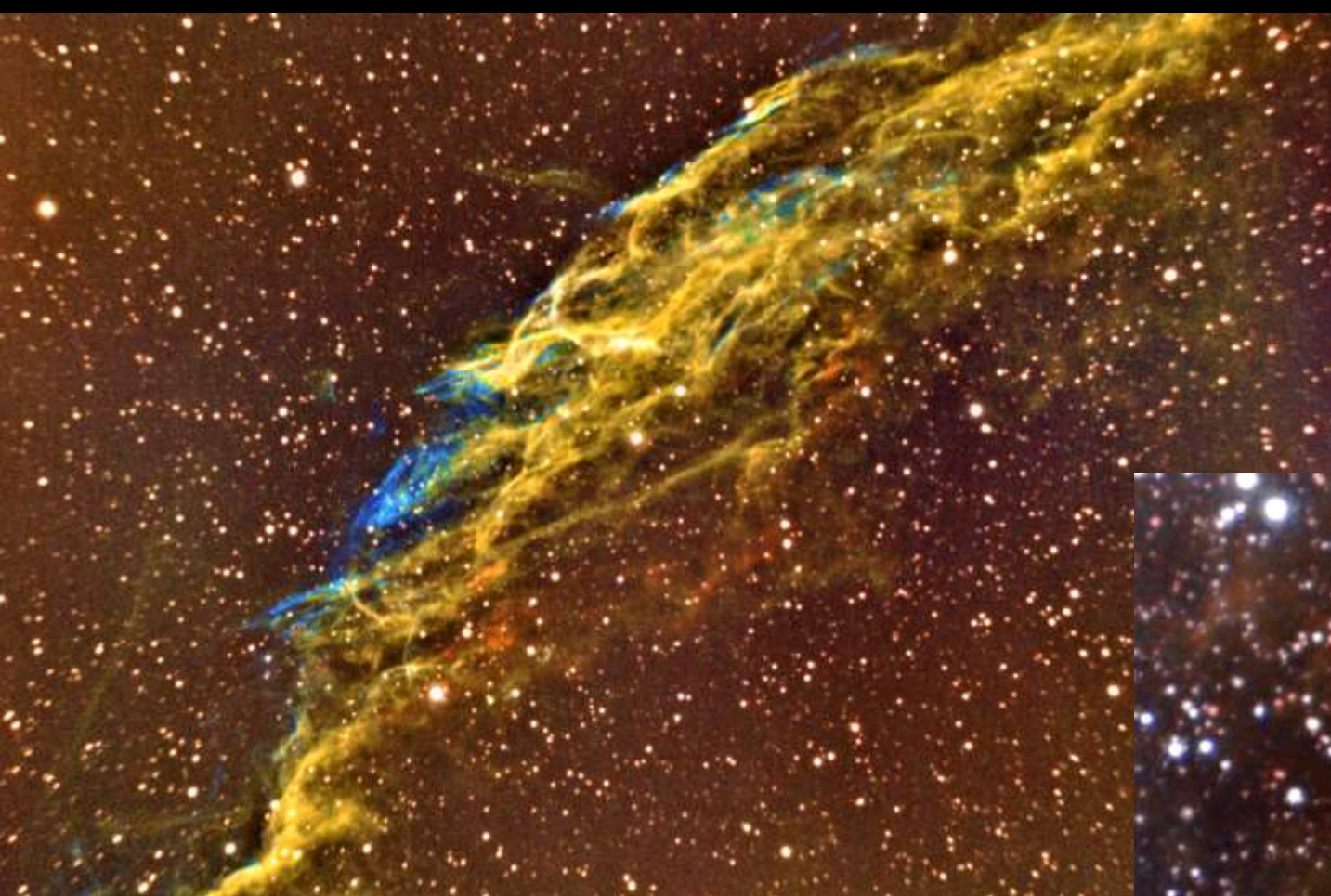
NGC 4631

M106



M42 Orion
RGB (4 h)





SII-H α -OIII (Hubble palette)

←NGC 6996 Cyg (8 h)

NGC 7635 Cas (6 h)



In prime focus: HyperStar → f/2.3 3° view

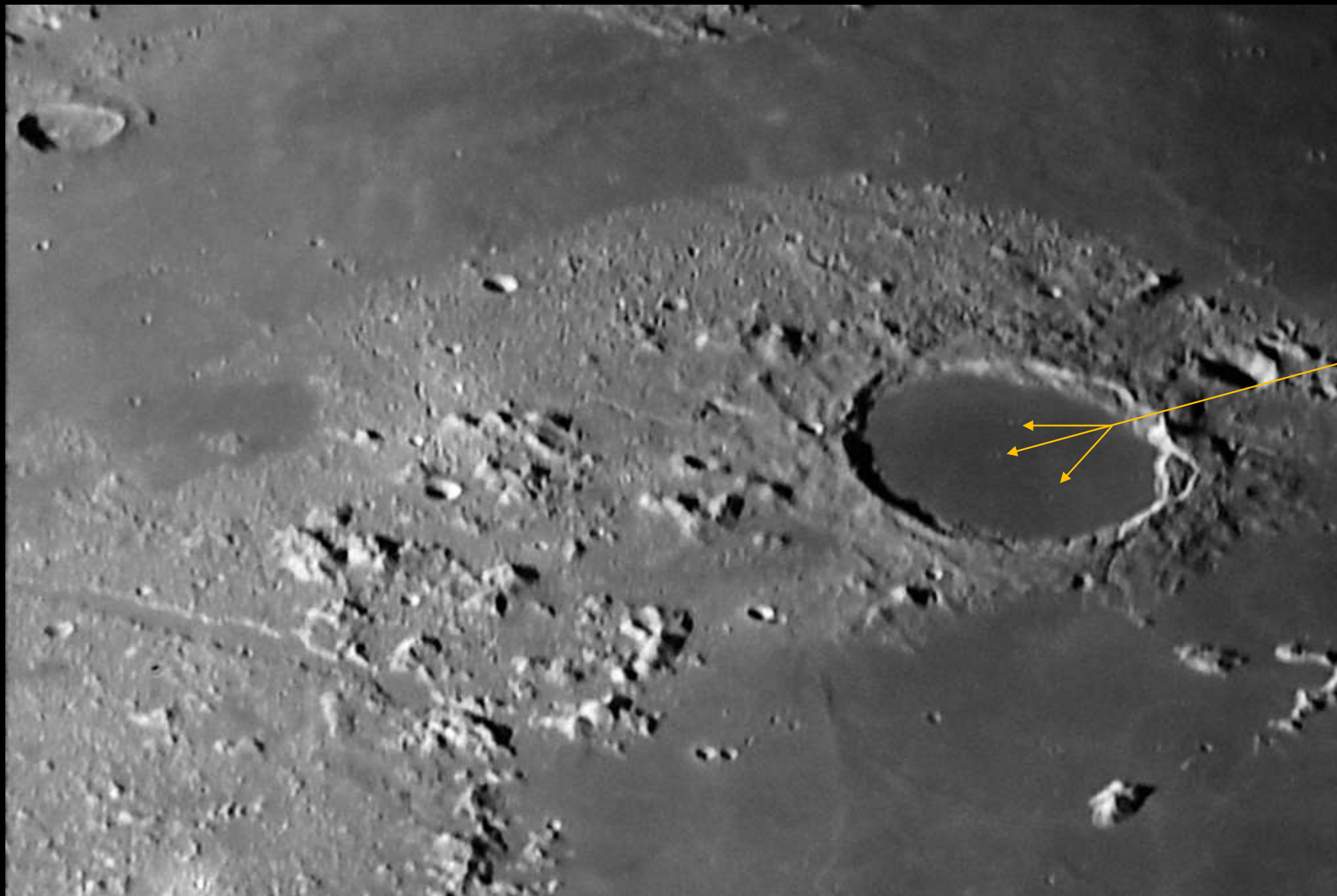


Hyperstar f/2.3
H α -OIII-SII
(10 h)



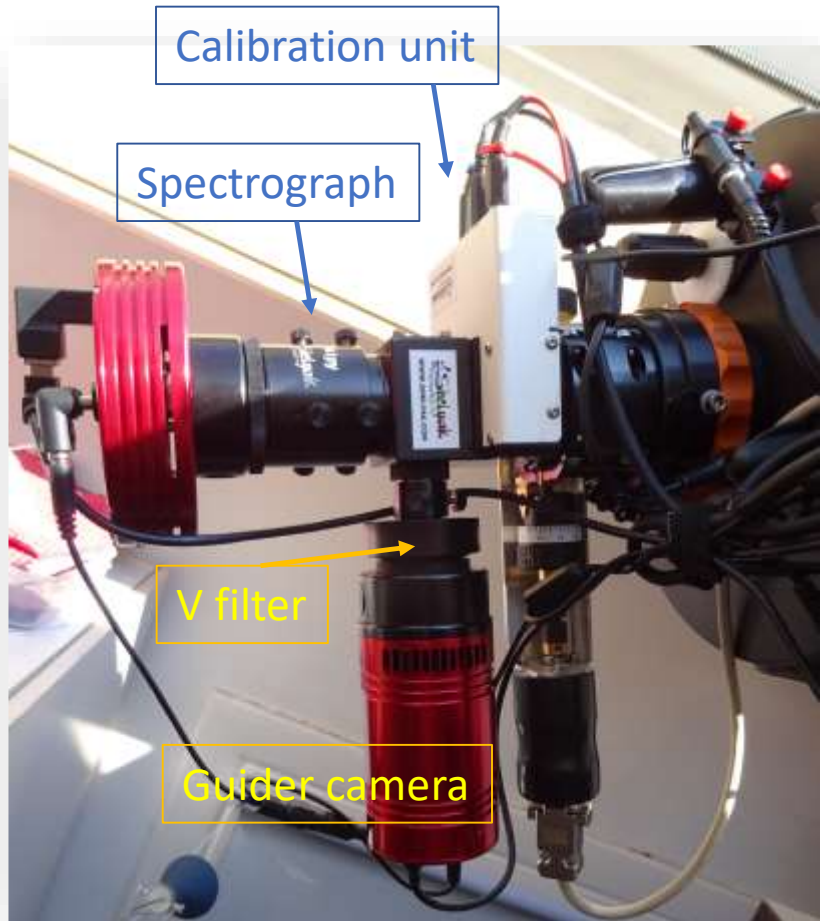
Rosetta
Hyperstar f/2.3
H α -OIII (11 h)



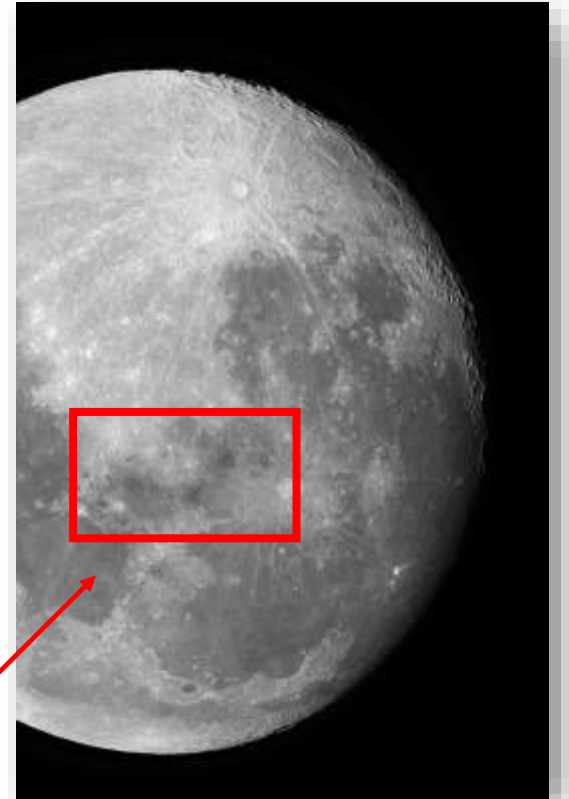


1 km

Spectrograph: Alpy (Shelyak)

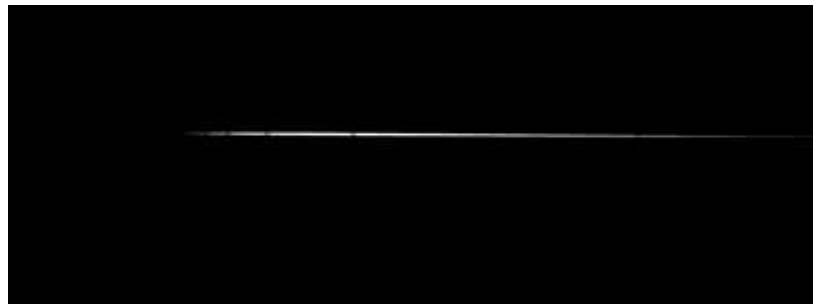
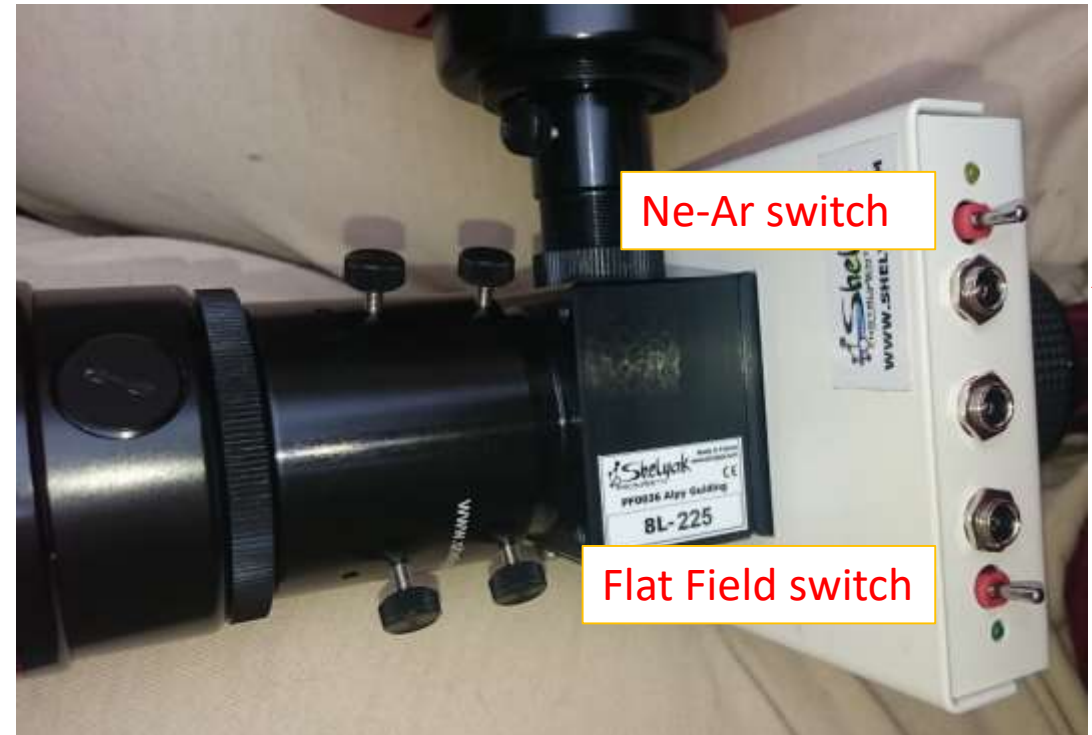
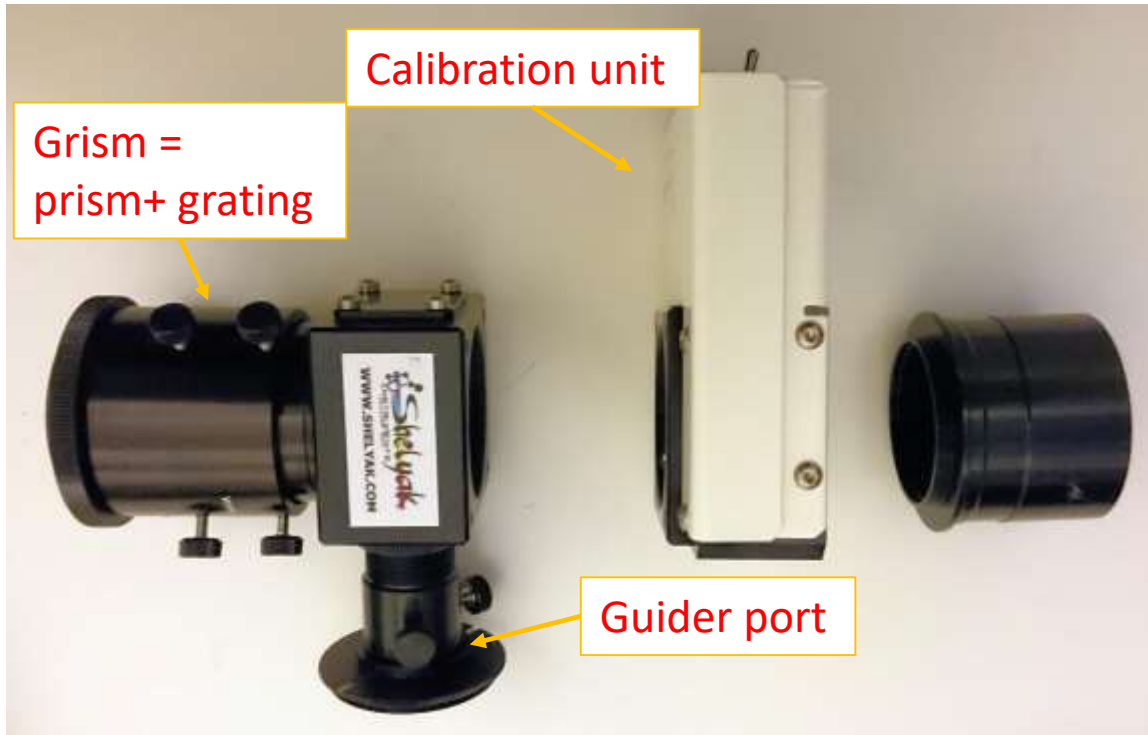


- Range: 3700 – 7500 Å
- Low resolution $\approx 400 - 600$
- Ne-Ar and FF calibration lamps
- Guiding module, slit 25 μm
- Guider camera, Field of View: 9' x 5'
- **Johnson V filter in light path of guider camera**
→ **Photometry**
- Acquisition software: *MaxImDL*



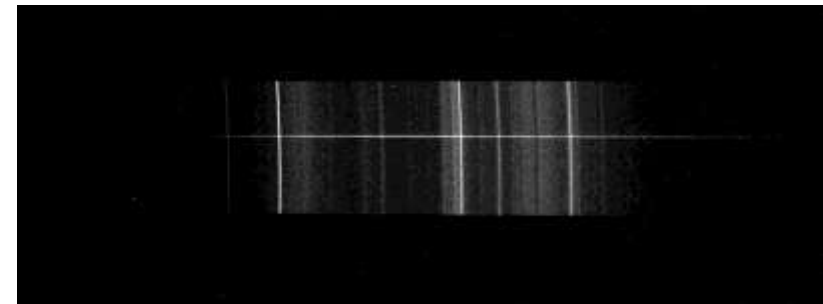
Shelyak Alpy spectrograph

Remote control

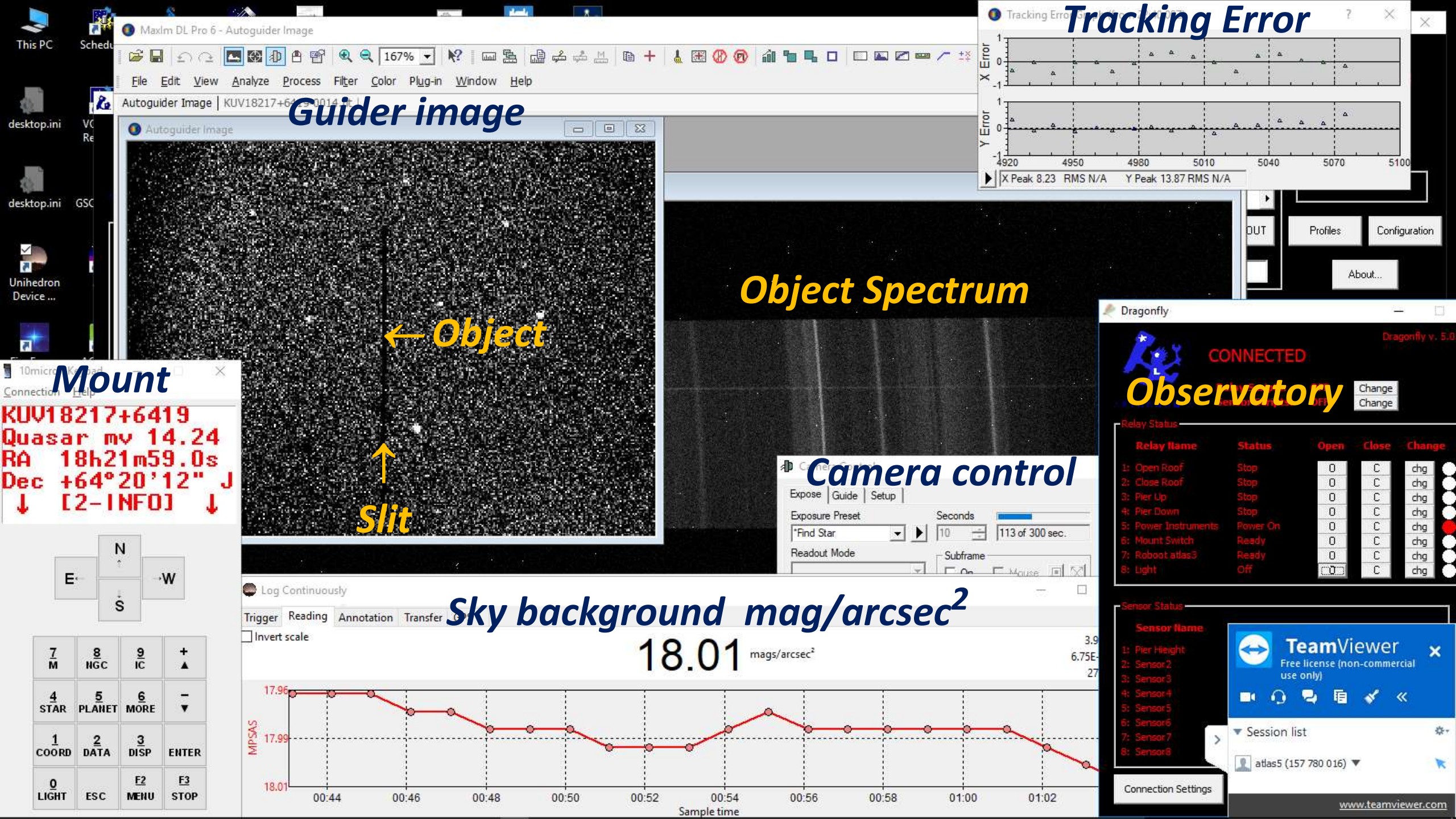


V= 5.9 Exp. 1 min

Binning
1x2



V = 14.6 Exp. 5 min



Guider image

← Object

↑ Slit

Object Spectrum

Camera control

Sky background mag/arcsec²

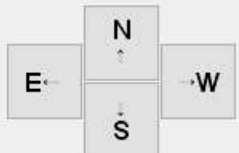
18.01 mags/arcsec²

Tracking Error

Observatory

Mount

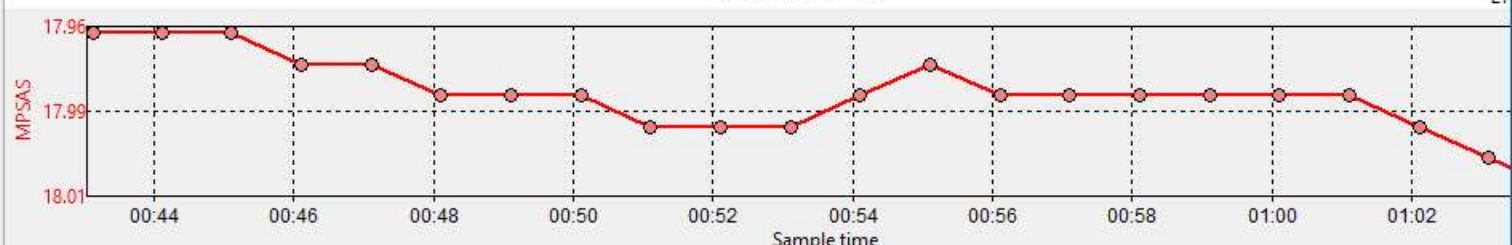
KUV18217+6419
Quasar mv 14.24
RA 18h21m59.0s
Dec +64°20'12" J
↓ [2-INFO] ↓



Log Continuously

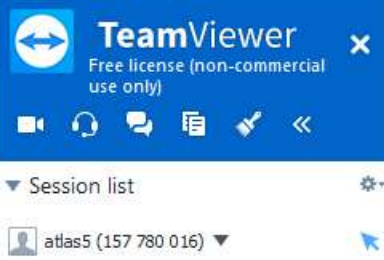
Trigger Reading Annotation Transfer

☐ Invert scale



Sensor Status

- Sensor Name
- 1: Pier Height
- 2: Sensor2
- 3: Sensor3
- 4: Sensor4
- 5: Sensor5
- 6: Sensor6
- 7: Sensor7
- 8: Sensor8



Connection Settings

www.teamviewer.com

Project examples with low-resolution spectroscopy

- Supernova
- Quasars
- Phaethon

SN2017eaw in NGC 6946 (Cepheus) “Fireworks galaxy”

13 May 2017

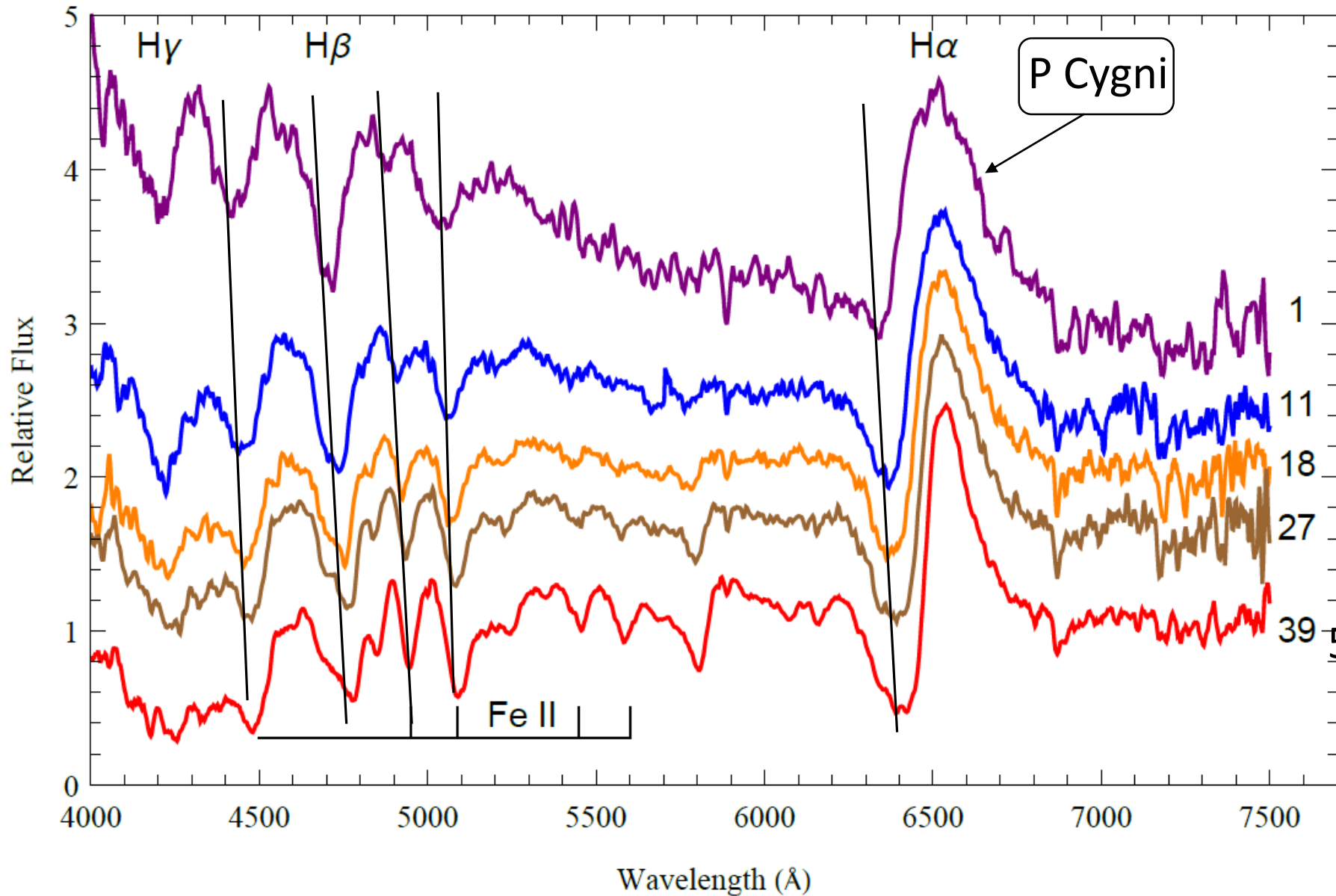
11 supernovae in last 100 yrs



8 spectra during 38 days

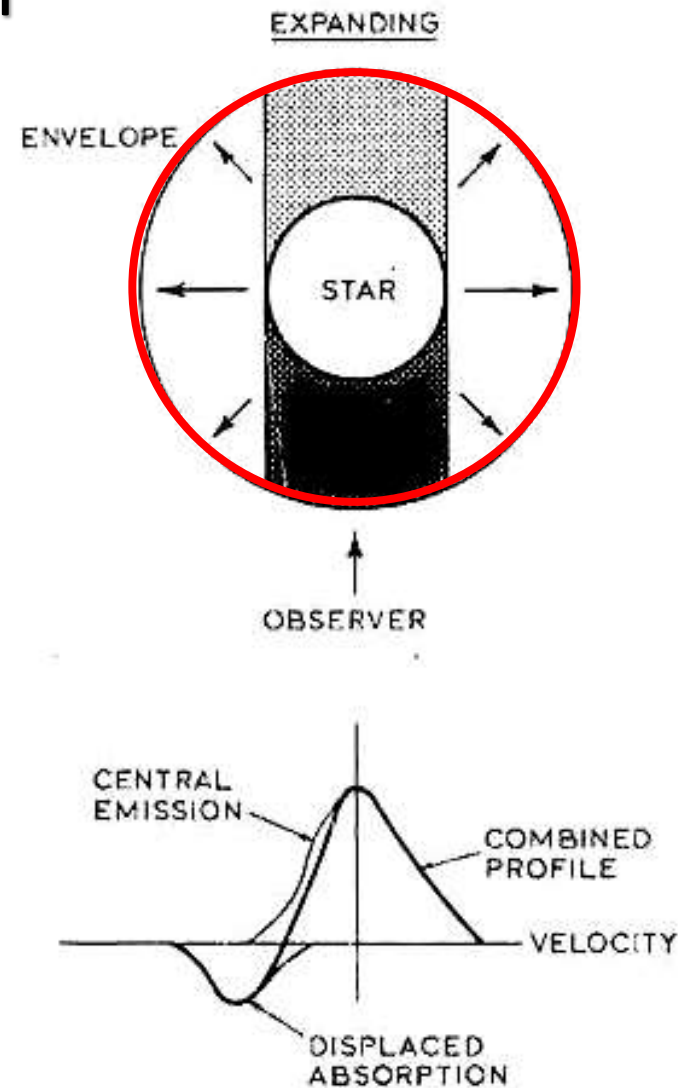
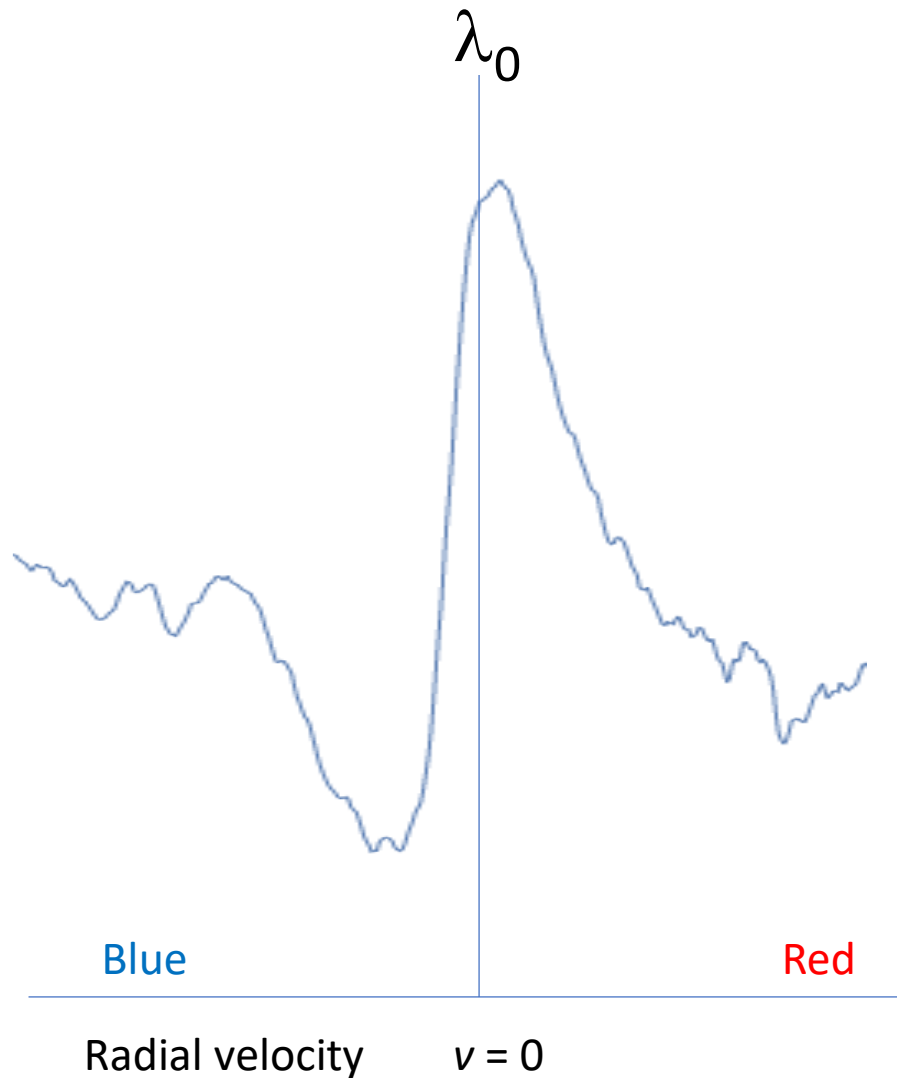
Starting 17 days after discovery

Spectral evolution SN2017eaw, 31 May 2017 – 8 July 2017

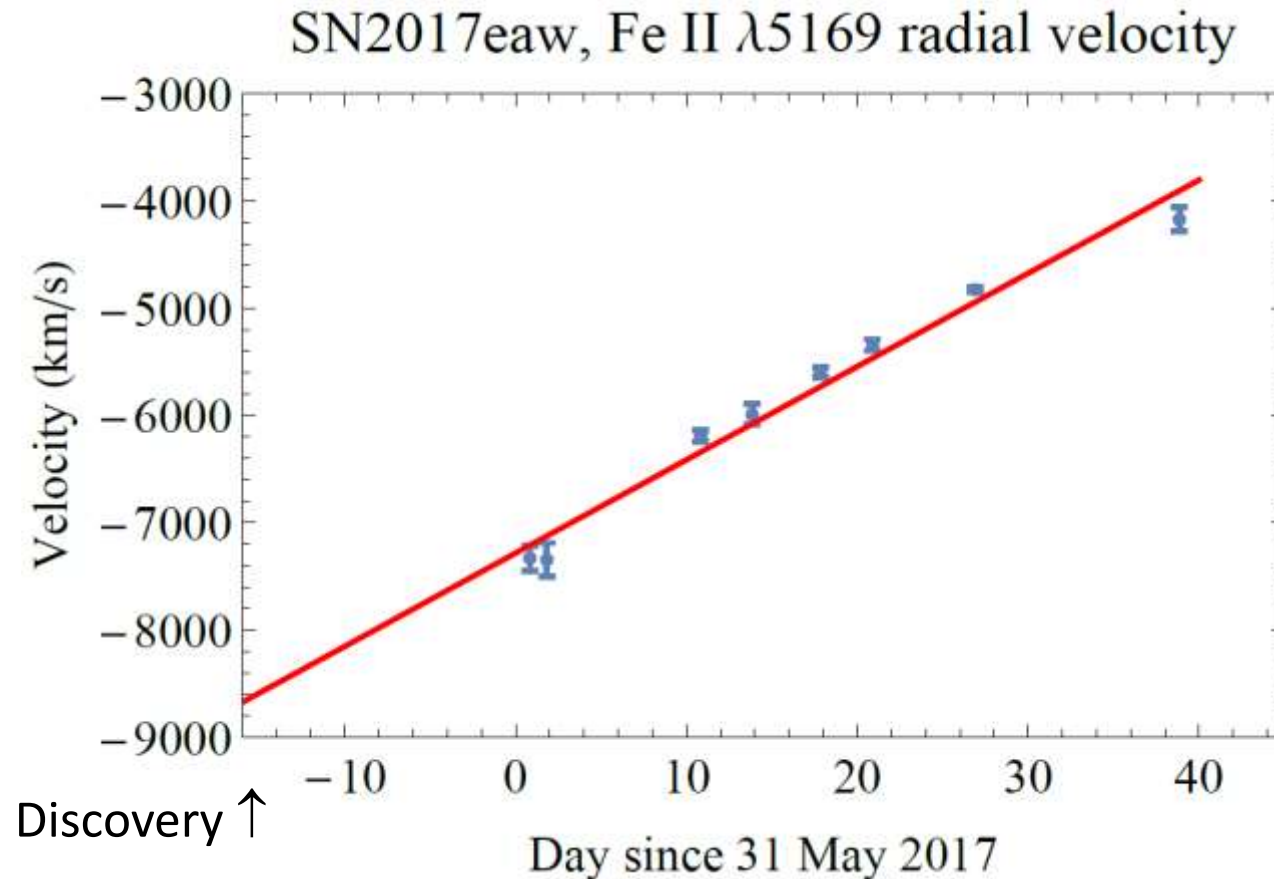
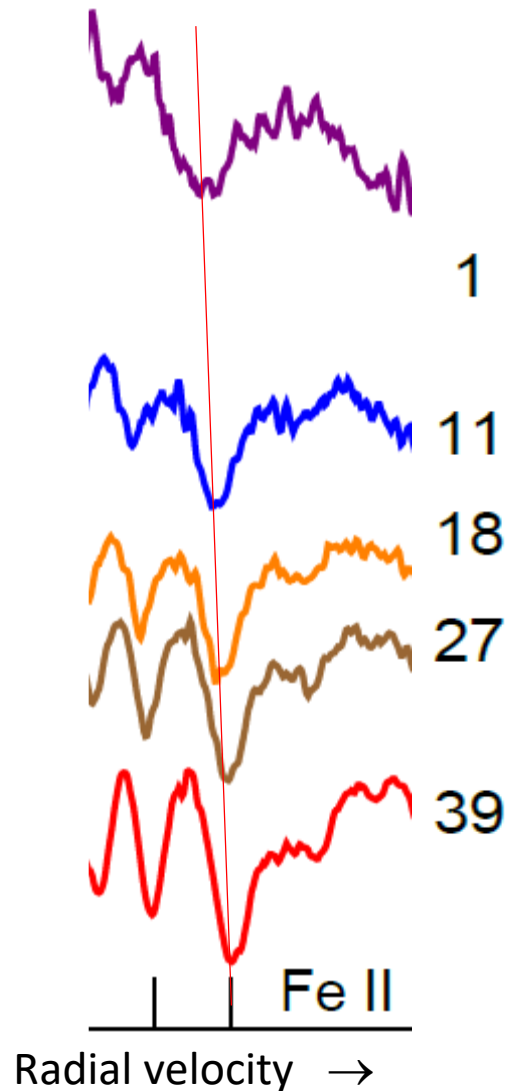


1. Color evolves from blue object to red
2. H lines → collapse of massive star into neutron star or black hole
3. Fe II lines grow stronger
4. H lines show P Cygni profile → outflow
5. Expansion velocity decreases

P Cygni profiles \Leftrightarrow expanding shell

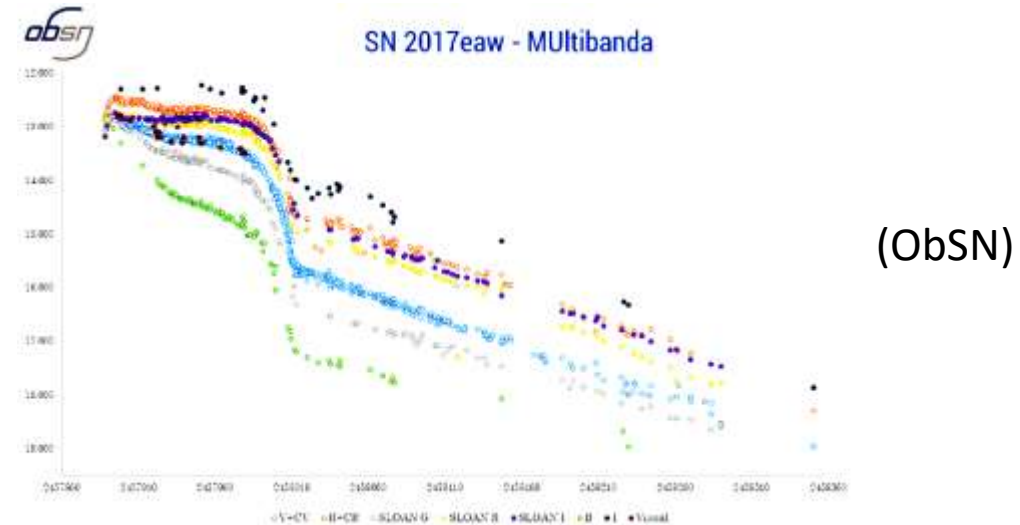
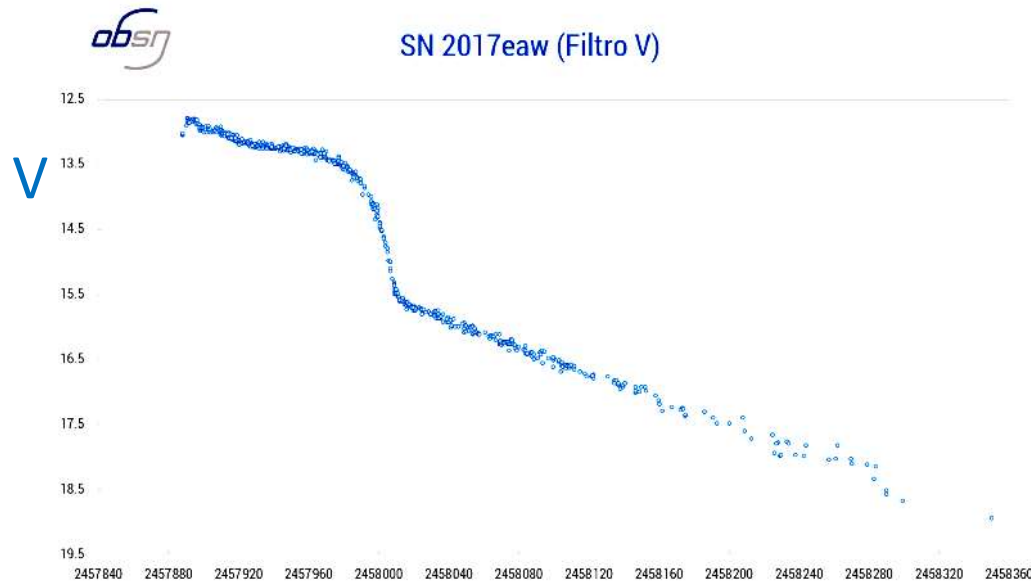


Decelerating velocity of expanding shell



Conclusion: supernova exploded with $v = 9000$ km/s

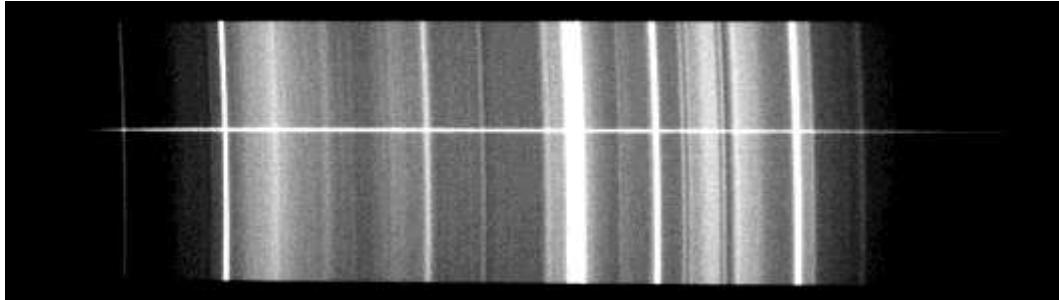
Supernova Type IIP (P → Plateau in light curve)



- (Type I: white dwarf of $1.4 M_{\odot}$ implodes to neutron star in a binary)
- Type II: Core collapse of massive star: 8 to $\sim 50 M_{\odot}$
- Remnant is neutron star ($1 M_{\odot}$) or black hole ($> 3 M_{\odot}$)
- In NGC 6946 all 11 supernovae were core-collapse SN (massive stars): never observed before; reason unknown

Sorry for C. Doppler (Salzburg, 1803)

Quasar 3C 273 (the easiest quasar in the sky, $V \approx 14.5$)

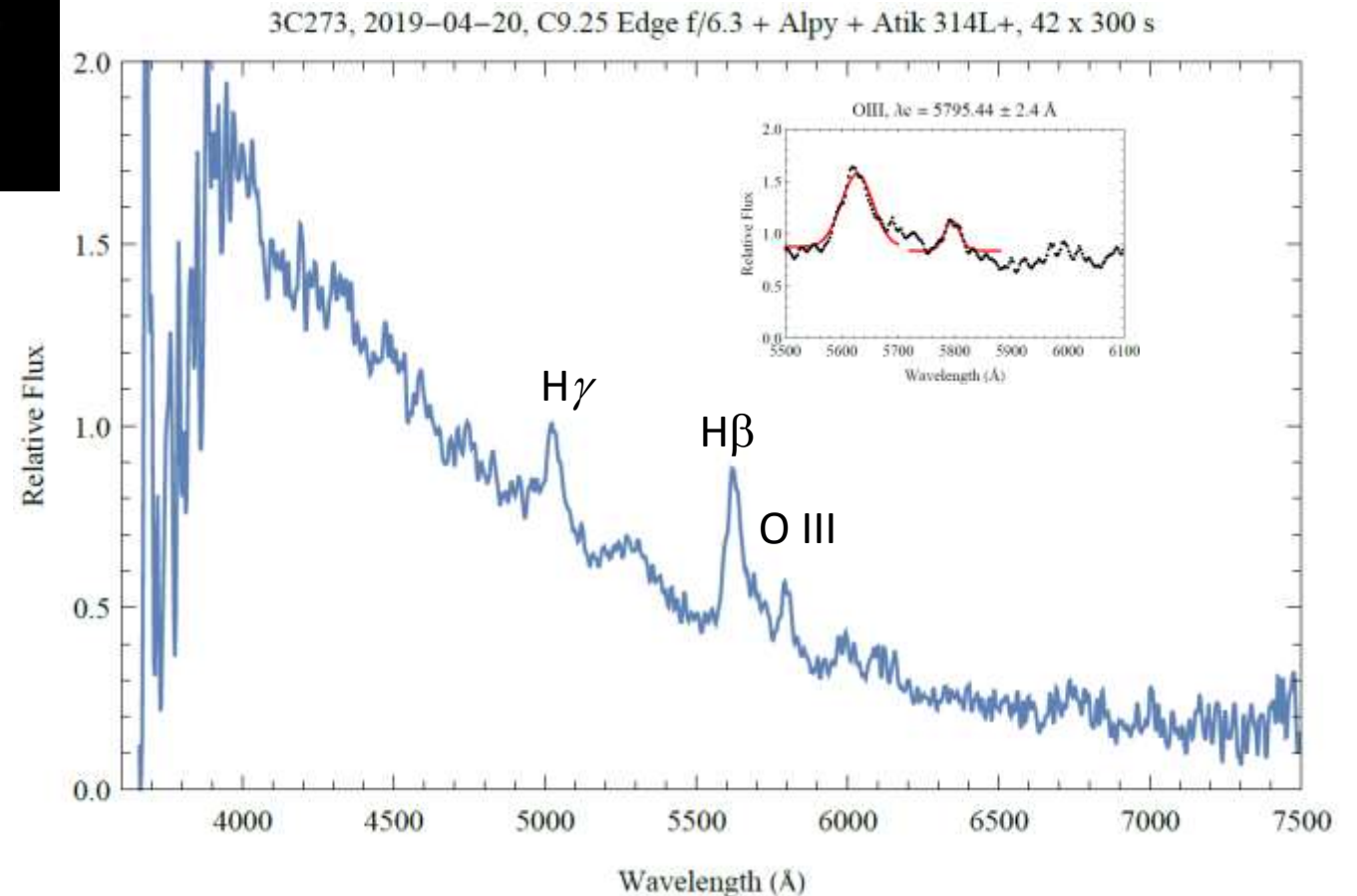


Redshift $z = \frac{\lambda - \lambda_0}{\lambda_0}$ *(not Doppler effect)*

Measured OIII, H β , H γ	SIMBAD
Redshift $z = 0.158$	$z = 0.158$
Velocity $v = 43650$ km/s	$v = 43750$ km/s
≈ 0.145 c	

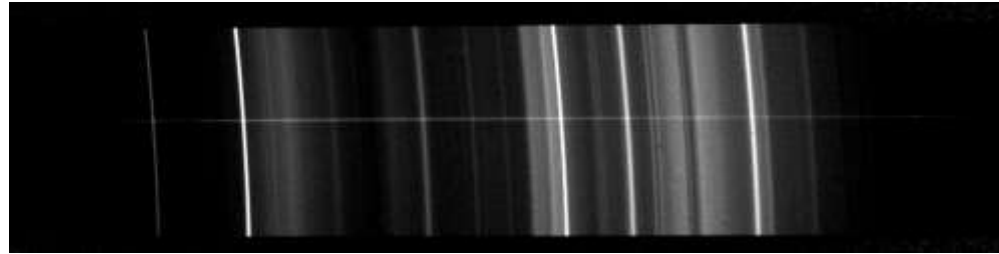
$z = 0.158$ means:
universe has expanded 1.158 larger

Look-back time is 2.0 Gyrs



Quasar HS0624+6907 in Cam, $t_{\text{exp}} = 5.2 \text{ h}$

HS0624+6907, 2019-03-29, C9.25 Edge f/6.3 + Alpy + Atik 314L+, 62 x 300 s



$$z = \frac{\lambda - \lambda_0}{\lambda_0}$$

Measured $H\beta$, $H\gamma$, $H\delta$

SIMBAD

Redshift $z = 0.37$

$z = 0.374$

Velocity $v = 91250 \text{ km/s}$

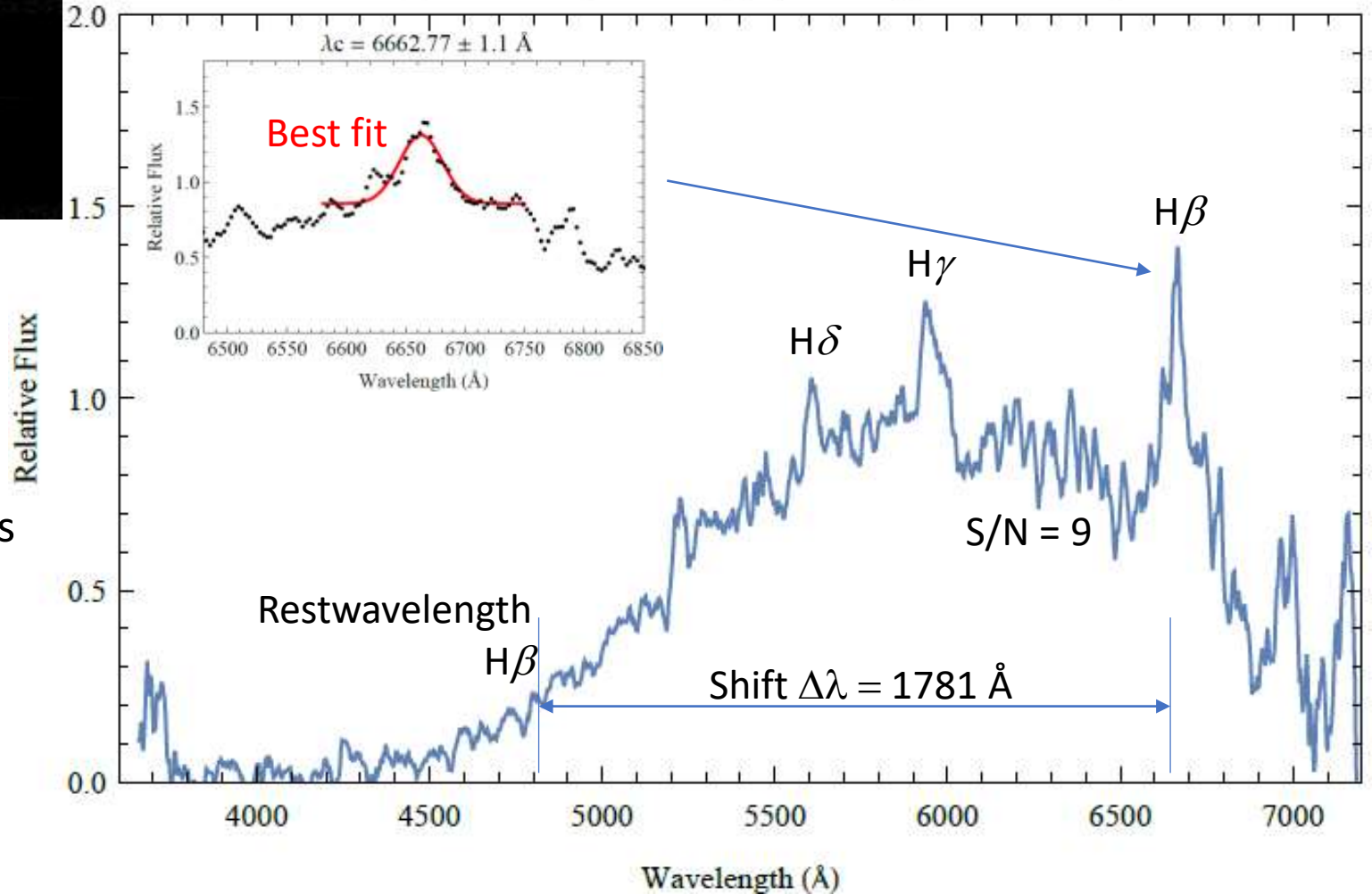
$v = 92171 \text{ km/s}$

$\approx 0.3 c$

$z = 0.37$ means:

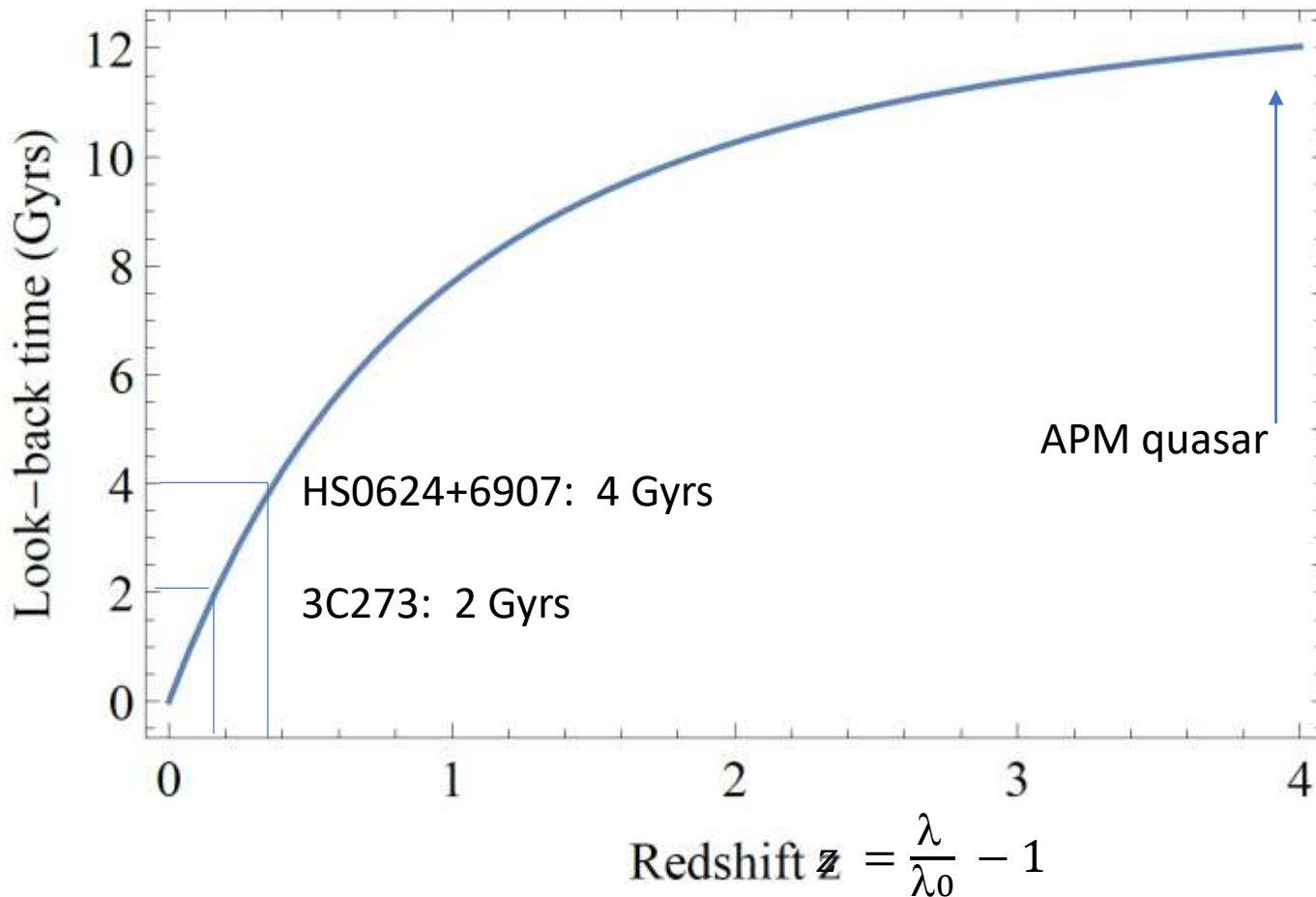
universe has expanded 1.37 larger

Look-back time is 4.0 Gyrs



Cosmology: look-back time:
depends on model of the universe
(Hubble constant, Flatness..)

$$\frac{1}{H_0} \int_0^z \frac{1}{(1+x) \sqrt{\Omega(1+x)^3 + \Omega_k(1+x)^2 + \Omega_\Lambda}} dx$$



Current model:

$$13.7 \text{ Gyrs} \int_0^z \frac{1}{(1+x) \sqrt{0.27(1+x)^3 + 0.73}} dx$$

Phaethon: closely passed the Earth in December 2017

- Who is Phaethon
- Why is Phaethon exceptional
- Simultaneous spectroscopy and photometry
- Scientific part: - rapid motion
 - brightness variations
 - spectroscopic identification

1. Who is Phaethon?

(Greek mythology, Ovidius)

Son of Helios

Allowed to drive the sun chariot

Lost control → collision course

Set the earth on fire

Killed by lightning strike from Zeus

Felt in river Eridanus

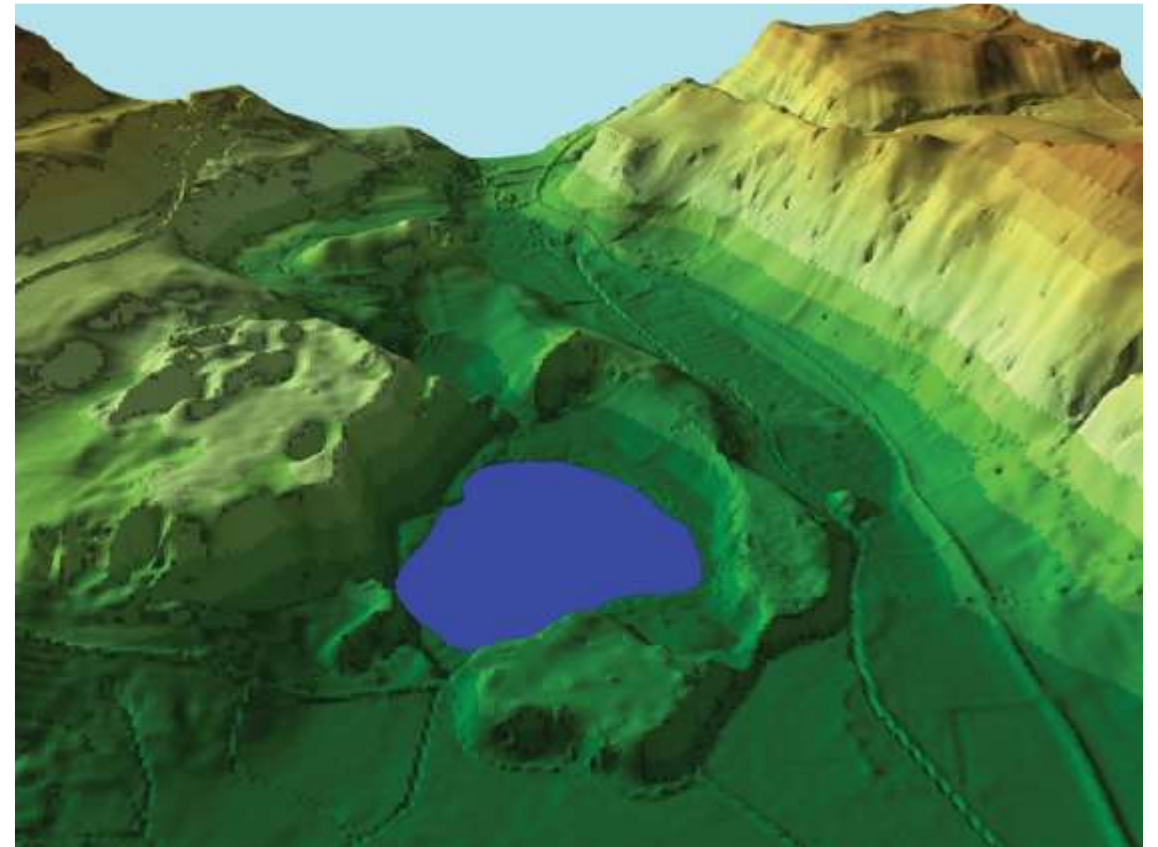
Tears from his sisters turned into amber

National Maritime Museum
Greenwich, London



Geomythology: origin of myth of Pheathon?

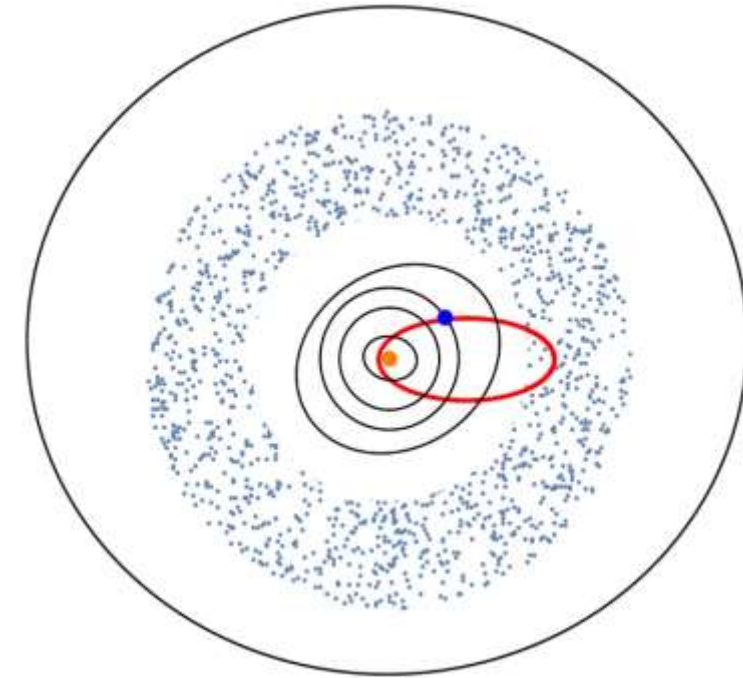
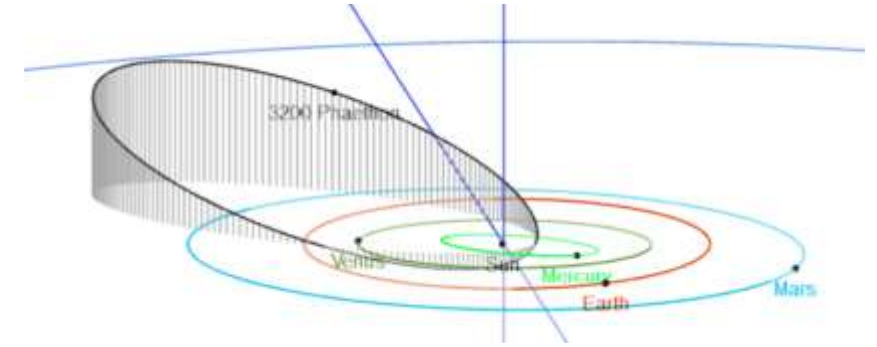
- Tüttensee crater, Bavaria
- Impact dated around Homer's time → given rise to myth



What makes Phaethon so exceptional

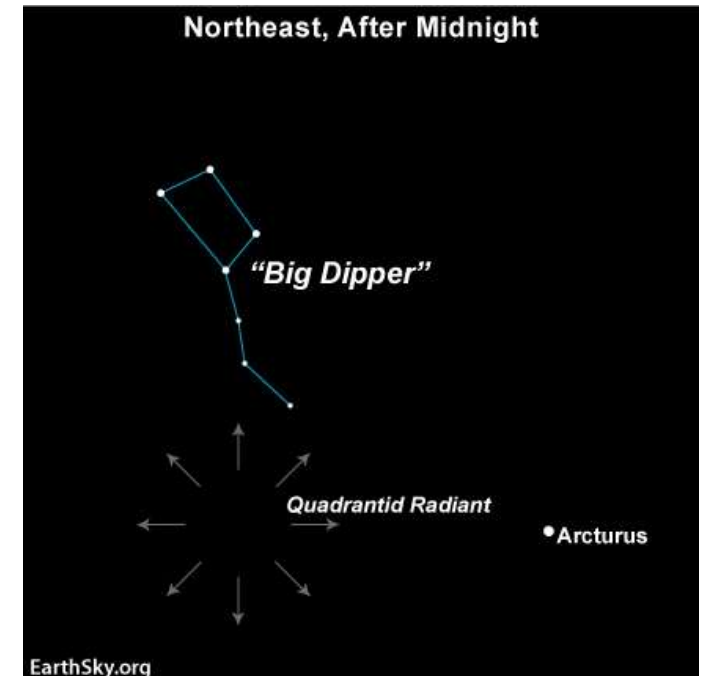
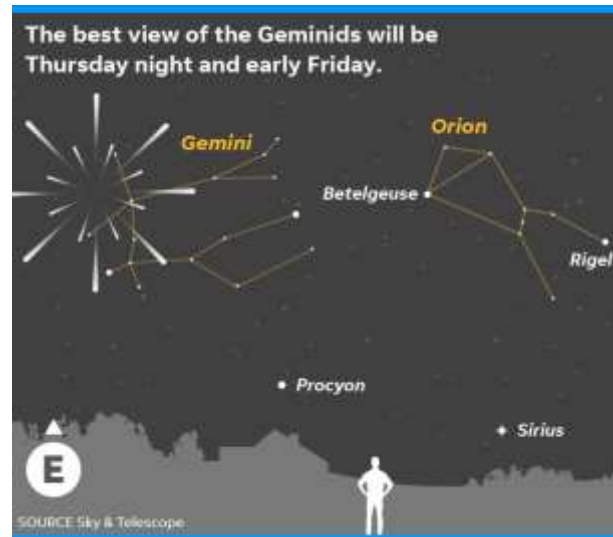
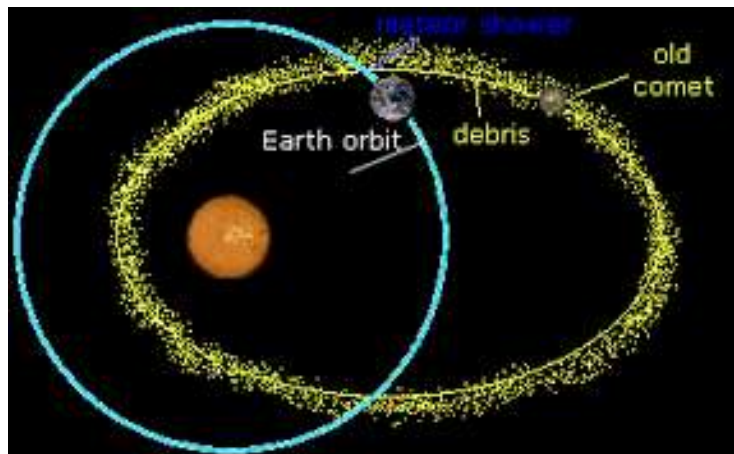
(discovered in 1983 by IRAS satellite)

- Orbital period = 1.43 years, very accurately known
 - Crosses Earth's orbit → **Potentially hazardous asteroid**
 - Eccentricity: 0.89 → **test for Einstein's General Relativity**
 - Parent body of the December ***Geminids*** meteor shower
- but:*** Temperature at perihelium: 600°C → **cannot be a comet**
- How come?**
- Closest approaches to Earth:
 - 16 December 2017 at 27 lunar distances
 - In 2093 at 8 lunar distances

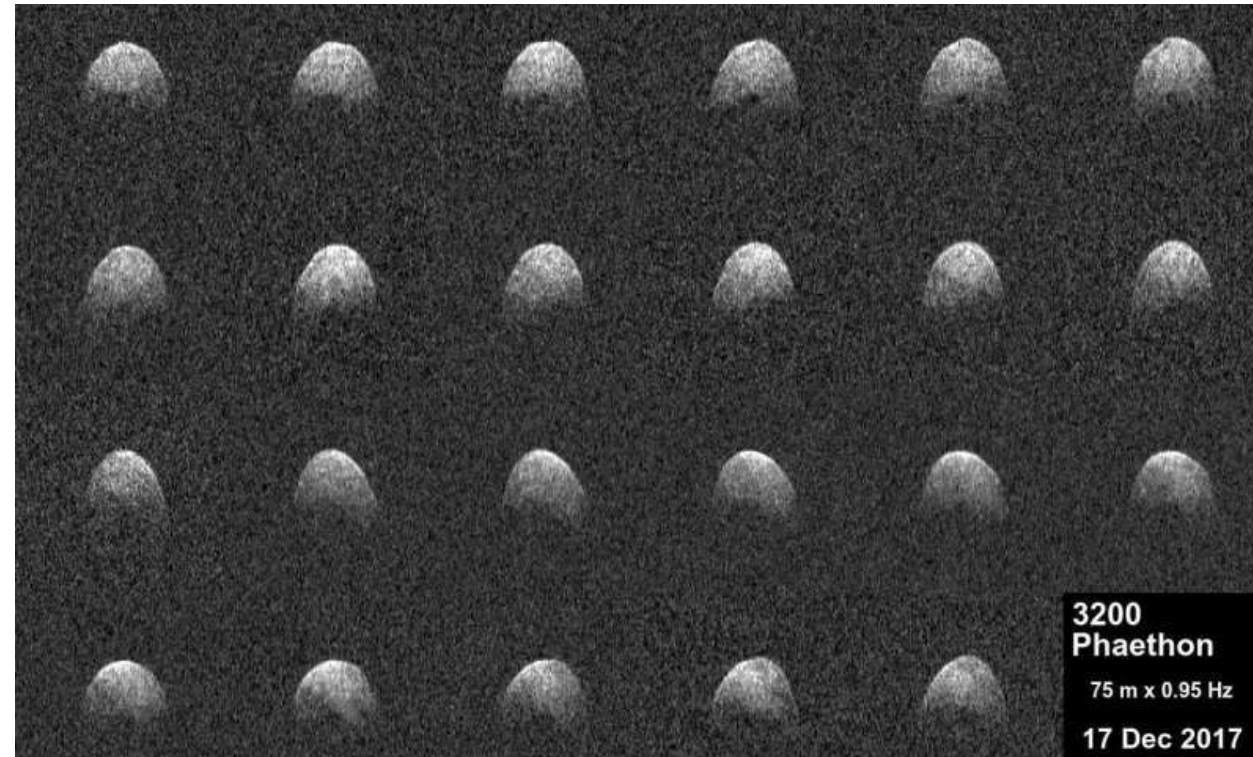


Meteor showers

- >100 well established (~ 800 known)
- All come from (older) comets
- Two exceptions: Geminids ← Phaethon
Quadrantides ← 2003 EH } How ??



Shape



<https://phys.org/news/2018-10-astrophysicists-asteroid-phaeton.html>



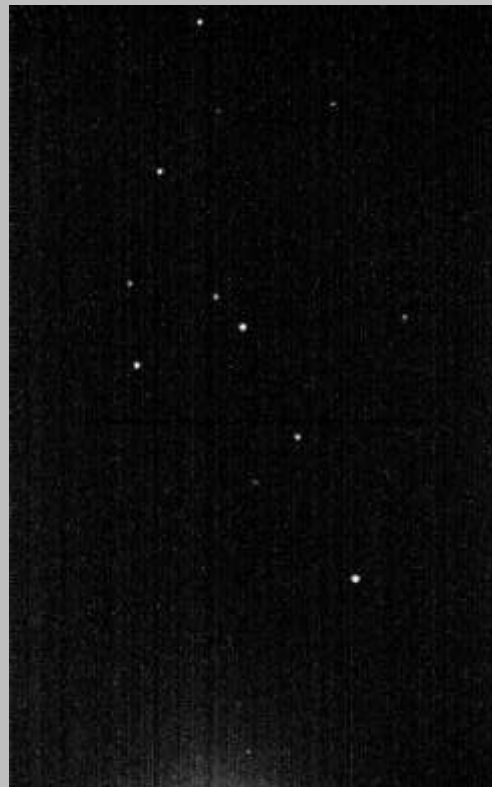
Model by S. de Vet

2. Imaging with guider camera

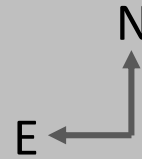
4 days before closest approach



17 min motion



200 x 5 sec exp.



2 hours before closest approach

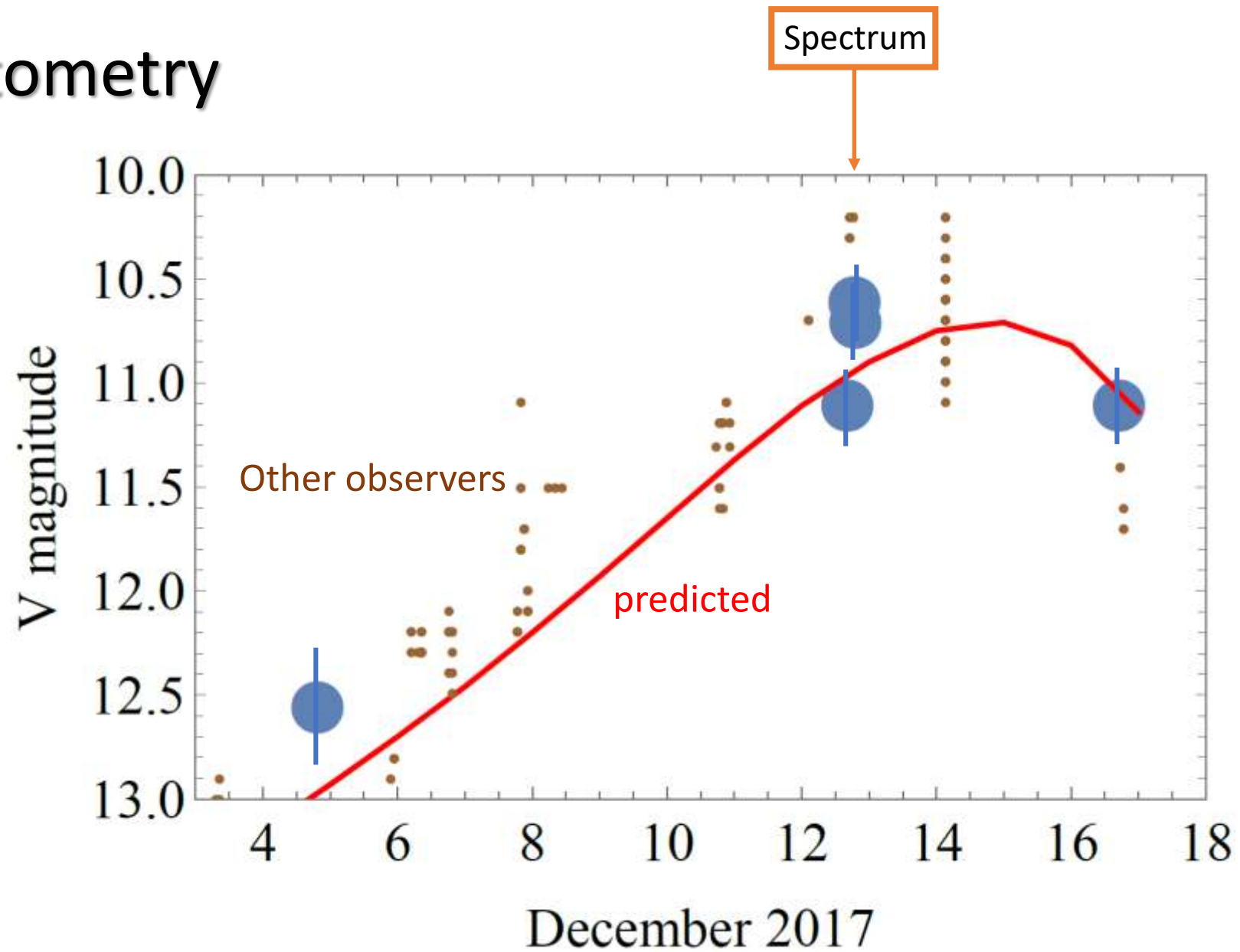


8 min star tracks
100 x 5 sec exp.

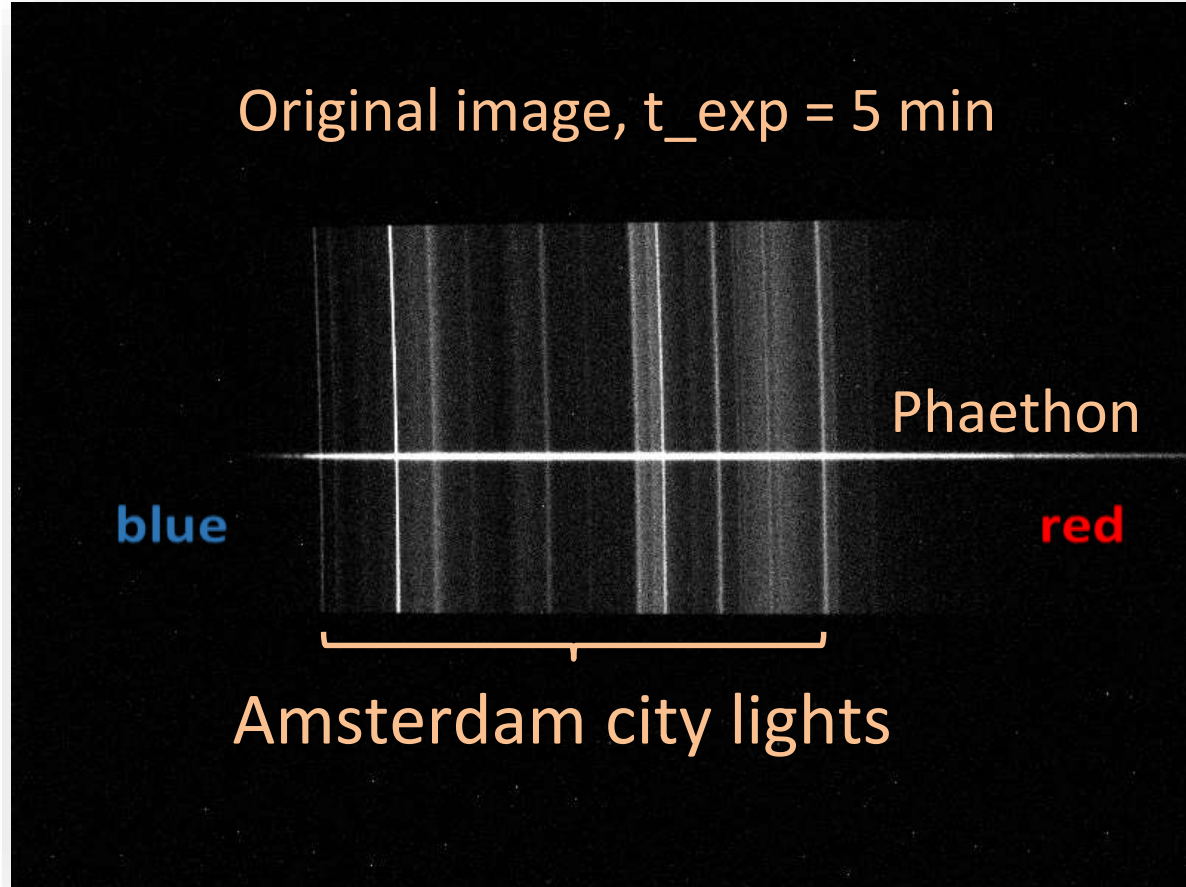


17 min motion
200 x 5 sec exp.

Photometry



Spectroscopy



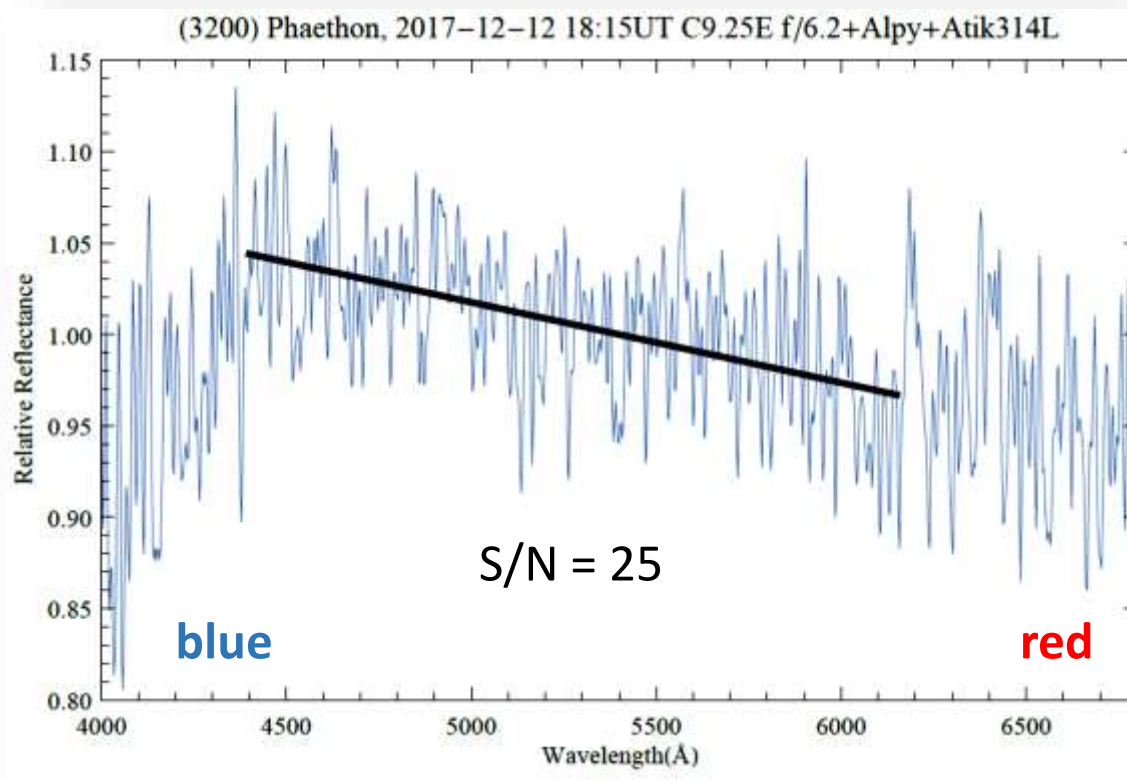
Processing (Mathematica):

Corrections for:

- Dark en Flatfield
- City-light background
- Spectrograph response curve

- Solar contribution removal with spectrum of HD 245, a solar-type star

Identification

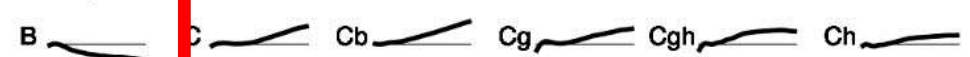


Bus-DeMeo Taxonomy Key

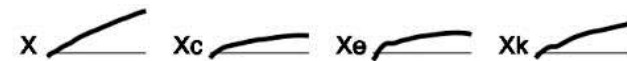
S-complex



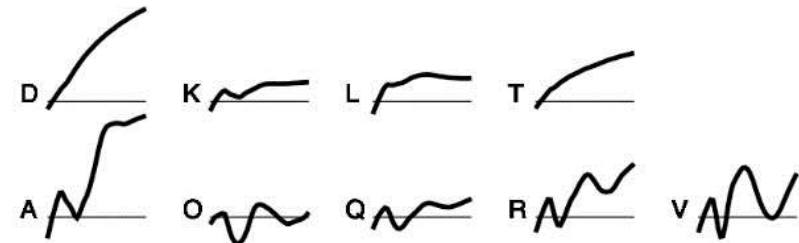
C-complex



X-complex



End Members



<http://smass.mit.edu/busdemeoclass.html>

F. E. DeMeo, R. P. Binzel, S. M. Slivan, and S. J. Bus. Icarus 202 (2009) 160-180

Phaethon reflection spectrum → C-complex Type B

(same as found by Licandro et al 2007, with the 4.2m William Herschel Telescope on La Palma, also with $t_{exp} = 5$ min)

Science conclusions and future

- Phaethon is a Carbon-type asteroid : **an exceptionally blue object**
 - hydrated-silicates surface, similar to Pallas-type meteorites
- During close passages near the Sun **‘active’ surface outbursts** occur (?)
 - fragmentation(?) → **Geminids meteor showers** (not understood)

- Future:
 - 2022 launch of Japanese DESTINY+ Flyby at 500 km
 - 14 December 2093 at 8 lunar distances (IAU 174 years)
 - Dynamical lifetime is few million years → orbit will desintegrate
 - Will Phaethon ever collide with the Earth ????

Greek mythology answer: no collision

